

Towards revised cloud radiation coupling for the COSMO Model

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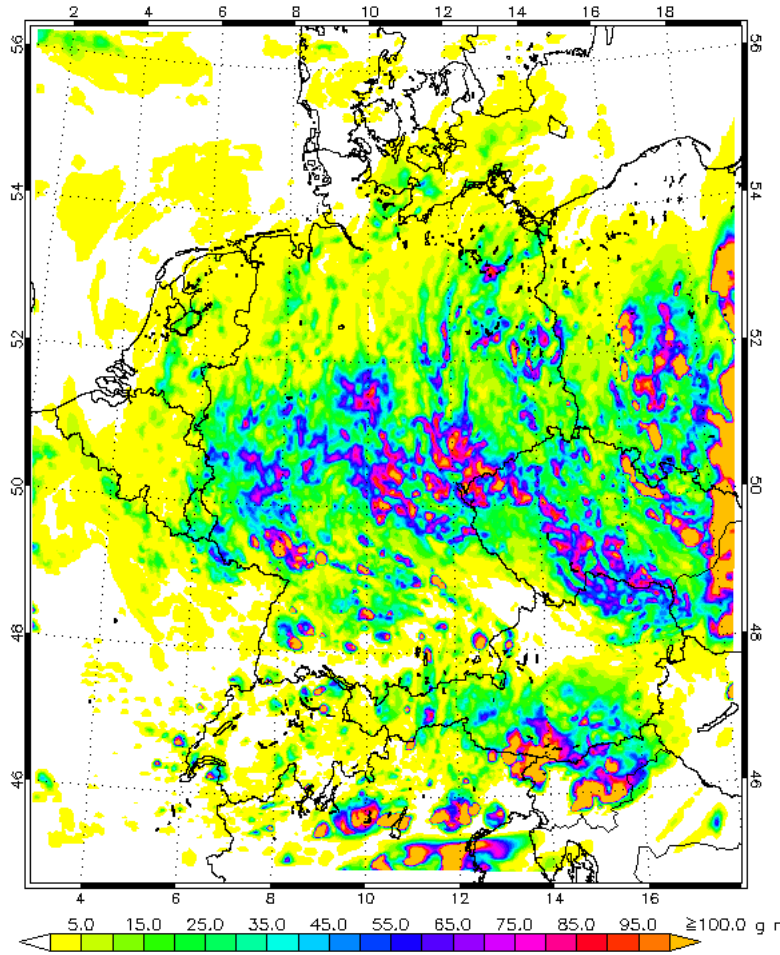
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²ETH Zurich

Motivation

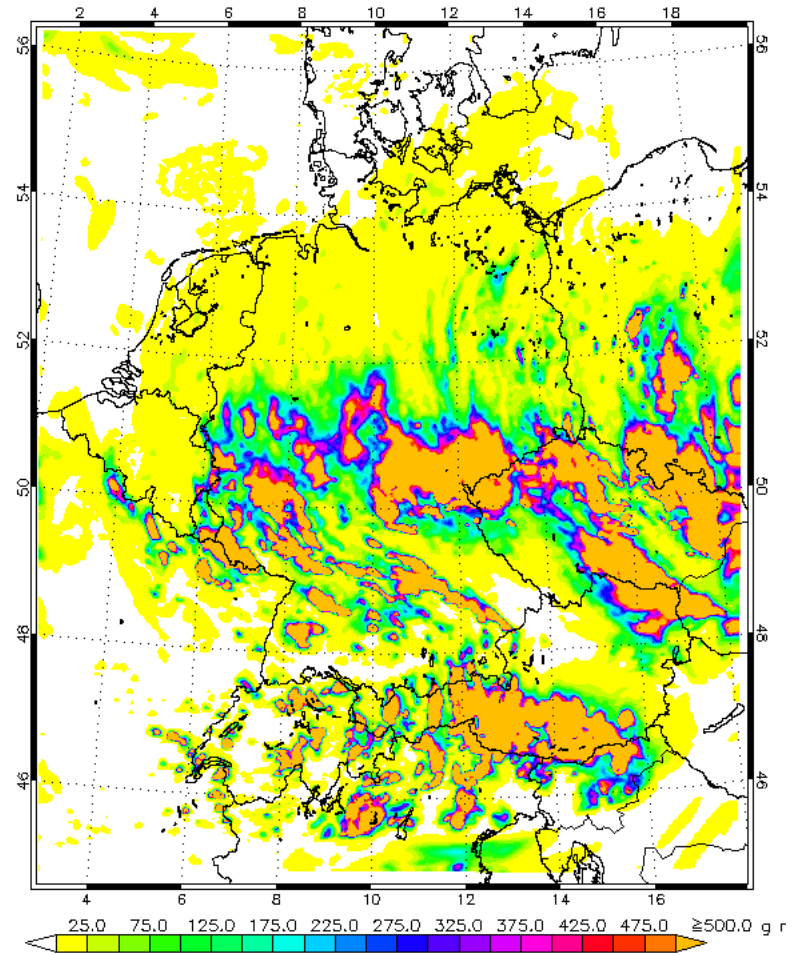
TQI 2012061300 UTC vv +00150000

TQS 2012061300 UTC vv +00150000



Colored field:	Min	Max	Mean	Var
	0.000	1314.700	14.906	988.132

gscp-4_radqfact-0.5_iradpar-1_aerosol-1

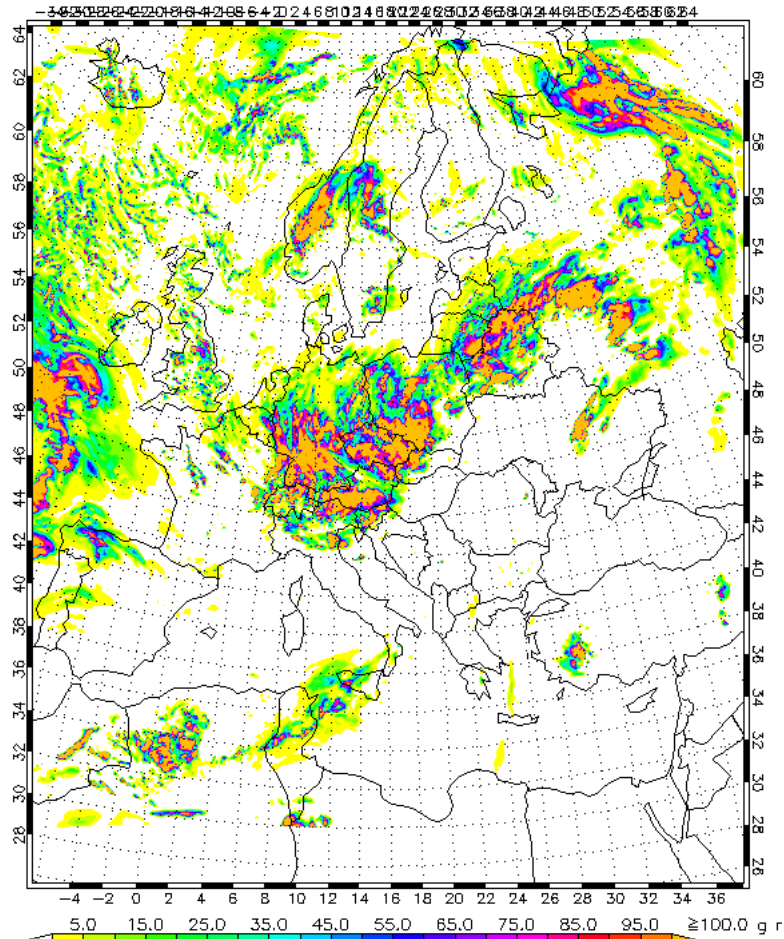


Colored field:	Min	Max	Mean	Var
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gscp-4_radqfact-0.5_iradpar-1_aerosol-1

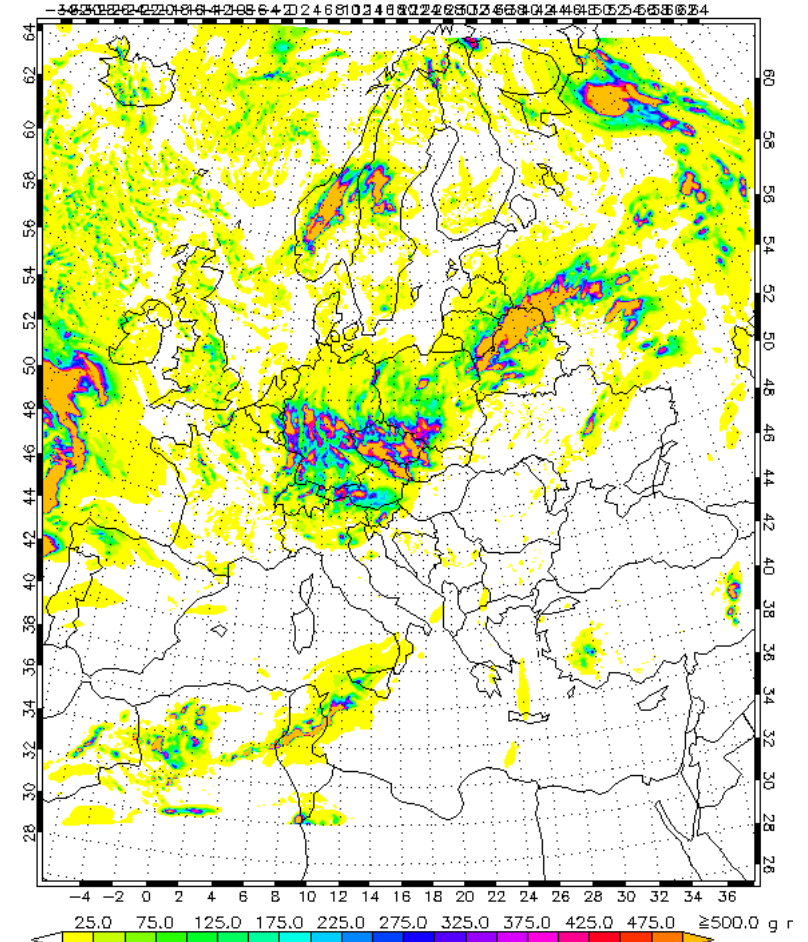
Motivation

TQI 2012061300 UTC vv +00150000



gscp-3_radqfact-0.5_iradpar-1_aerosol-1_NOLANDMASK

TQS 2012061300 UTC vv +00150000



gscp-3_radqfact-0.5_iradpar-1_aerosol-1_NOLANDMASK

Some issues in cloud-radiation coupling

- In the COSMO radiation scheme (Ritter & Geleyn 1992), the optical properties (extinction coeff., single scattering albedo, asymmetry factor) only depend on q_c respectively q_i .
 - Nowadays more modern parameterizations based on an effective radius R_e are available.
 - From inherent assumptions about $N(D)$ and particle shapes in state-of-the-art microphysical models R_e can be deduced.
 - Then: optical properties = fct (q_x, R_e)
- For the gridscale clouds, only a fraction $k=0.5$ of q_i and/or q_c is considered in the radiation scheme („tentative“ effective factor to take into account subgrid scale variability) ρ_c
- Radiation scheme is only aware of cloud droplets and cloud ice!

1) Revised optical properties

→ $\sigma_{\text{ext}}, \omega, g = \text{fct}(R_e, q_x, \lambda)$

→ For cloud droplets:

Parameterisation of Hu and Stamnes (1993), spectrally remapped to the 8 spectral intervals of RG92 (3 visible, 5 infrared).

→ For cloud ice:

→ Visible spectral region: optionally

Key et al. (2002), assuming horizontally aligned hexagonal plates or
Fu et al. (1998), assuming randomly oriented hexagonal needles

→ Infrared region:

Fu et al. (1996), assuming randomly oriented hexagonal needles

1a) Cloud droplets (Hu & Stamnes)

→ If grid scale $q_c > 0$: from cloud microphysics:

$$f(D) = N_0 D^\mu e^{-\lambda D}$$

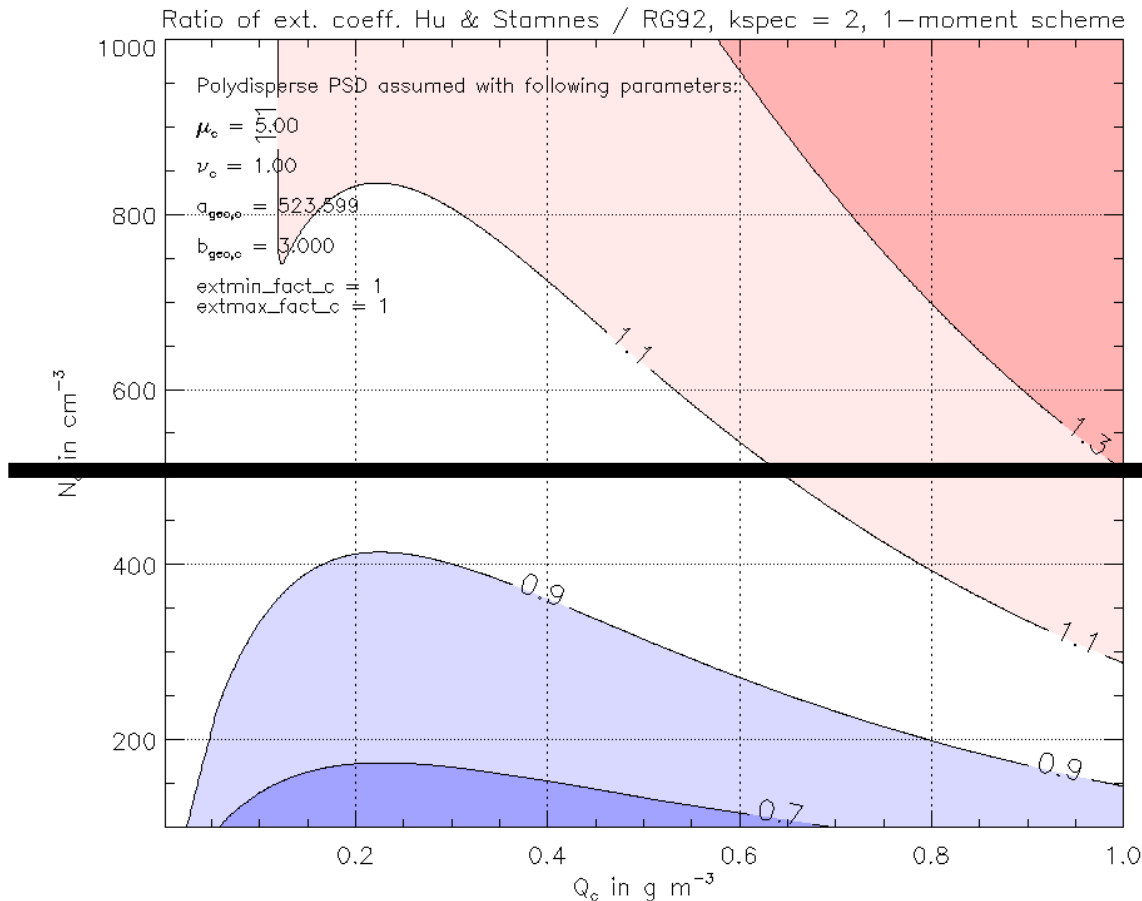
$$\mu = 5.0$$

$$N_c = \text{cloud_num}$$

q_c prognostic

Spectral interval „2“
(visible range)

σ_{ext} ratio HS / RG92



1a) Cloud droplets (Hu & Stamnes)

→ If grid scale $q_c > 0$: from cloud microphysics:

$$f(D) = N_0 D^\mu e^{-\lambda D}$$

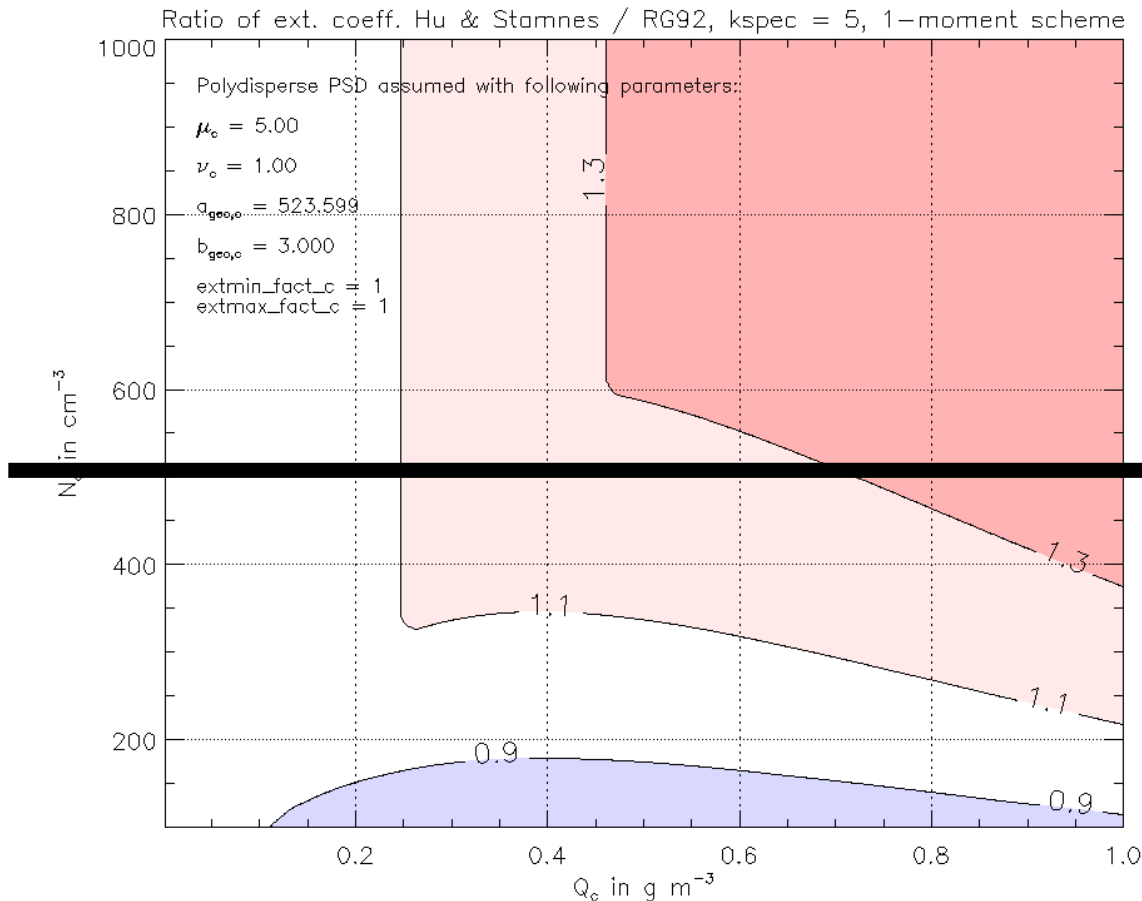
$$\mu = 5.0$$

$$N_c = \text{cloud_num}$$

q_c prognostic

Spectral interval „5“
(infrared range)

σ_{ext} ratio HS / RG92



1a) Cloud droplets (Hu & Stammes)

Deutscher Wetterdienst
Wetter und Klima aus einer Hand



→ Pure subgrid scale clouds ??? → $R_e = 10 \mu\text{m}$



1b) Cloud ice (visible; Fu et al.)

→ If grid scale $q_i > 0$: from cloud microphysics:

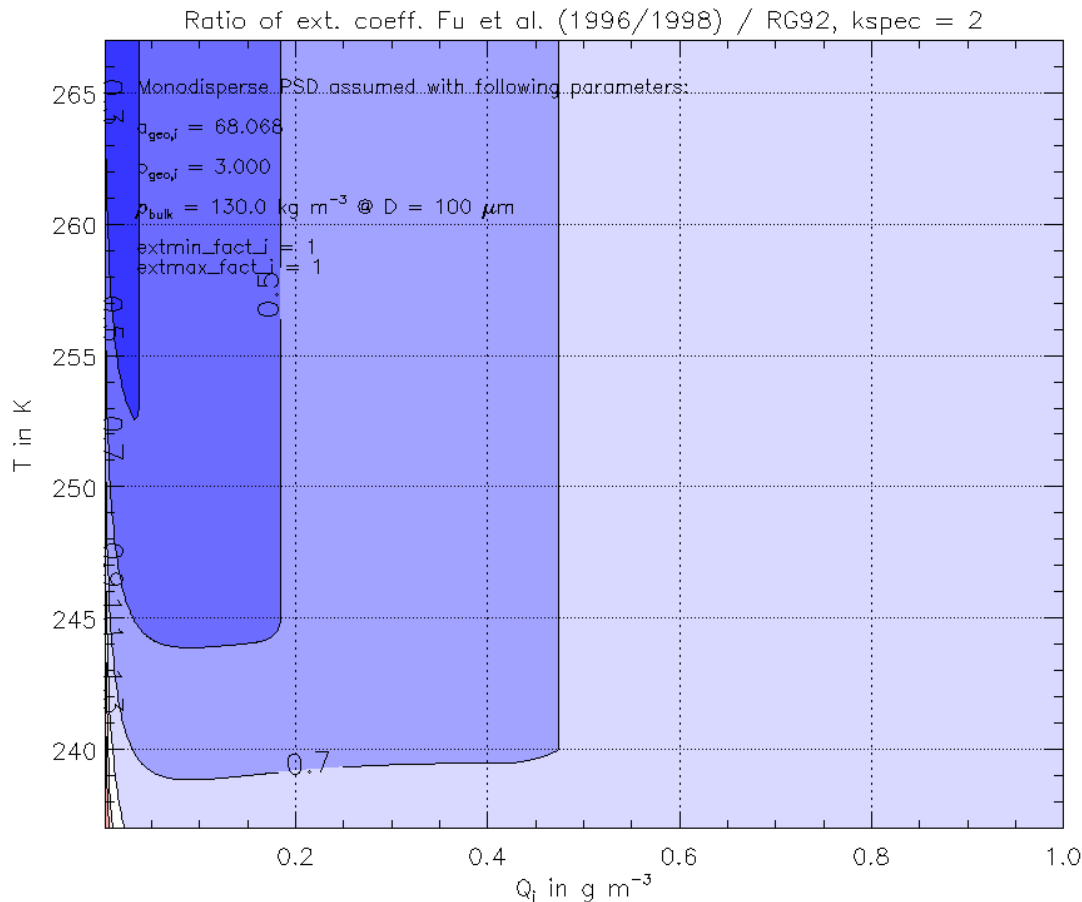
$$f(D) = \text{monodispers}$$
$$N_i(T) = a \exp(b(T_3 - T))$$

q_i prognostic

$$m_i = 130 D^3 \text{ (SI-units)}$$

Spectral interval „2“
(visible range)

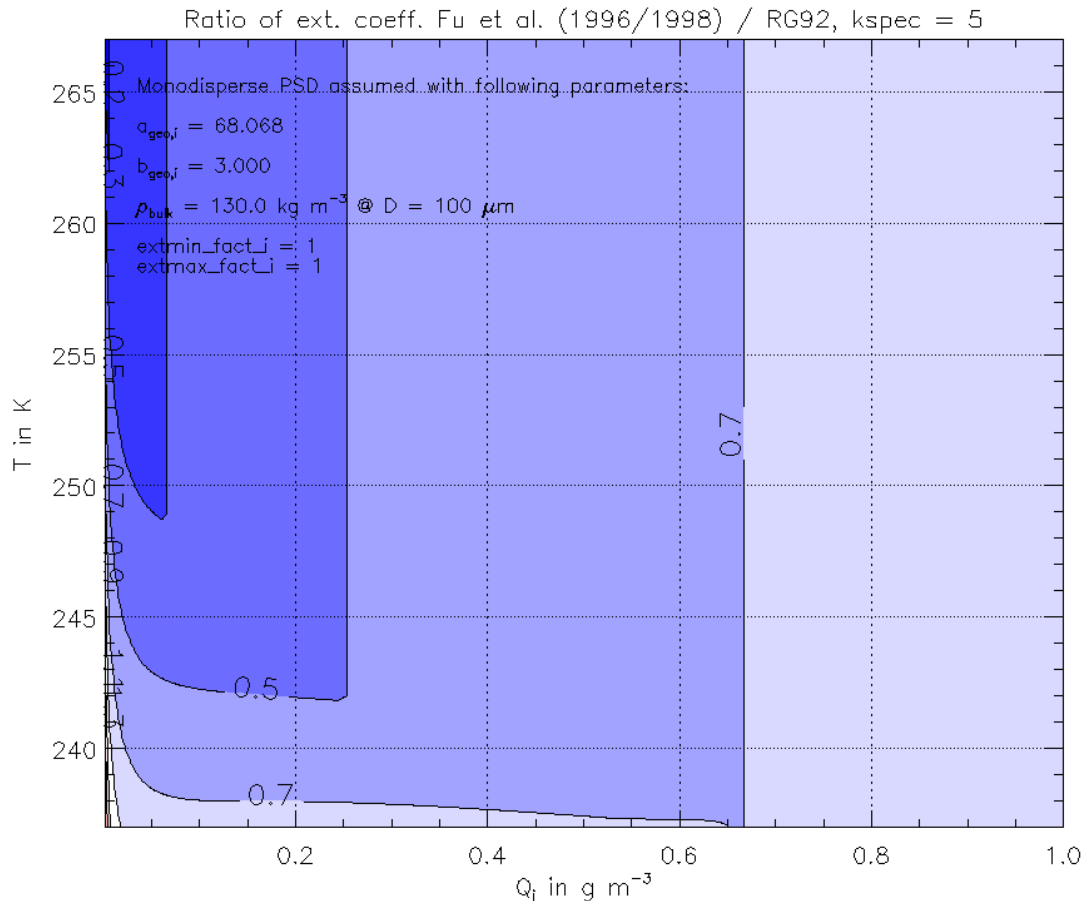
σ_{ext} ratio Fu / RG92



1b) Cloud ice (infrared; Fu et al.)

→ If grid scale $q_i > 0$: from cloud microphysics:

$f(D)$ = monodispers
 $N_i(T) = a \exp(b(T_3 - T))$
 q_i prognostic
 $m_i = 130 D^3$ (SI-units)



Spectral interval „5“
(infrared range)

σ_{ext} ratio Fu / RG92

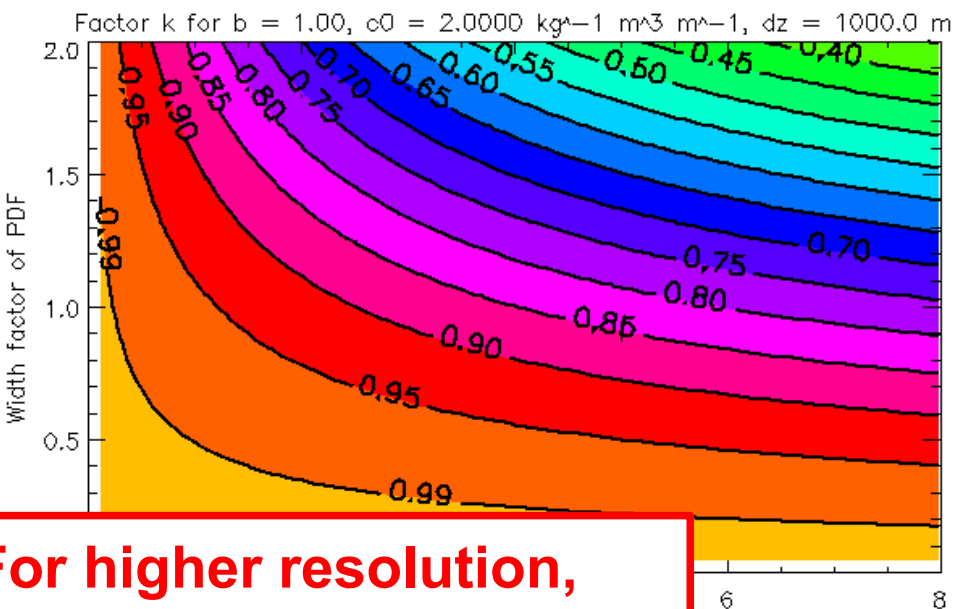
1b) Cloud ice (infrared; Fu et al.)

→ Pure subgrid scale ice clouds ??? → $R_e = 20 \mu\text{m}$

2) Subgridscale variability, factor k

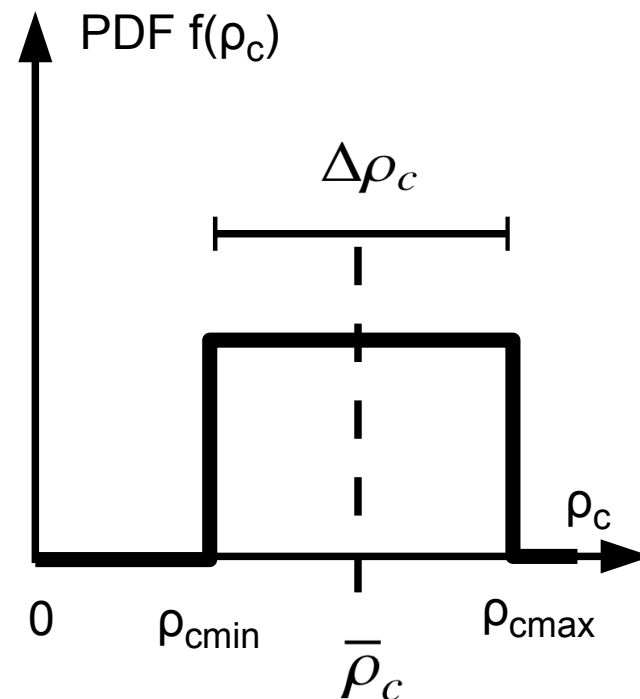
$$e^{(-k c_0 \bar{\rho}_c \Delta z)} \stackrel{!}{=} \int_0^{\infty} PDF(\rho_c) e^{(-c_0 \rho_c \Delta z)} d\rho_c$$

$$k = \frac{1}{\bar{\tau} c_0 \bar{\rho}_c \Delta z} \ln \left[\frac{\Delta \rho_c c_0 \Delta z}{\exp(-c_0 \rho_{cmin} \Delta z) - \exp(-c_0 \rho_{cmax} \Delta z)} \right]$$



**For higher resolution,
k = 0.5 seems too low!**

cloudy grid box ρ_c



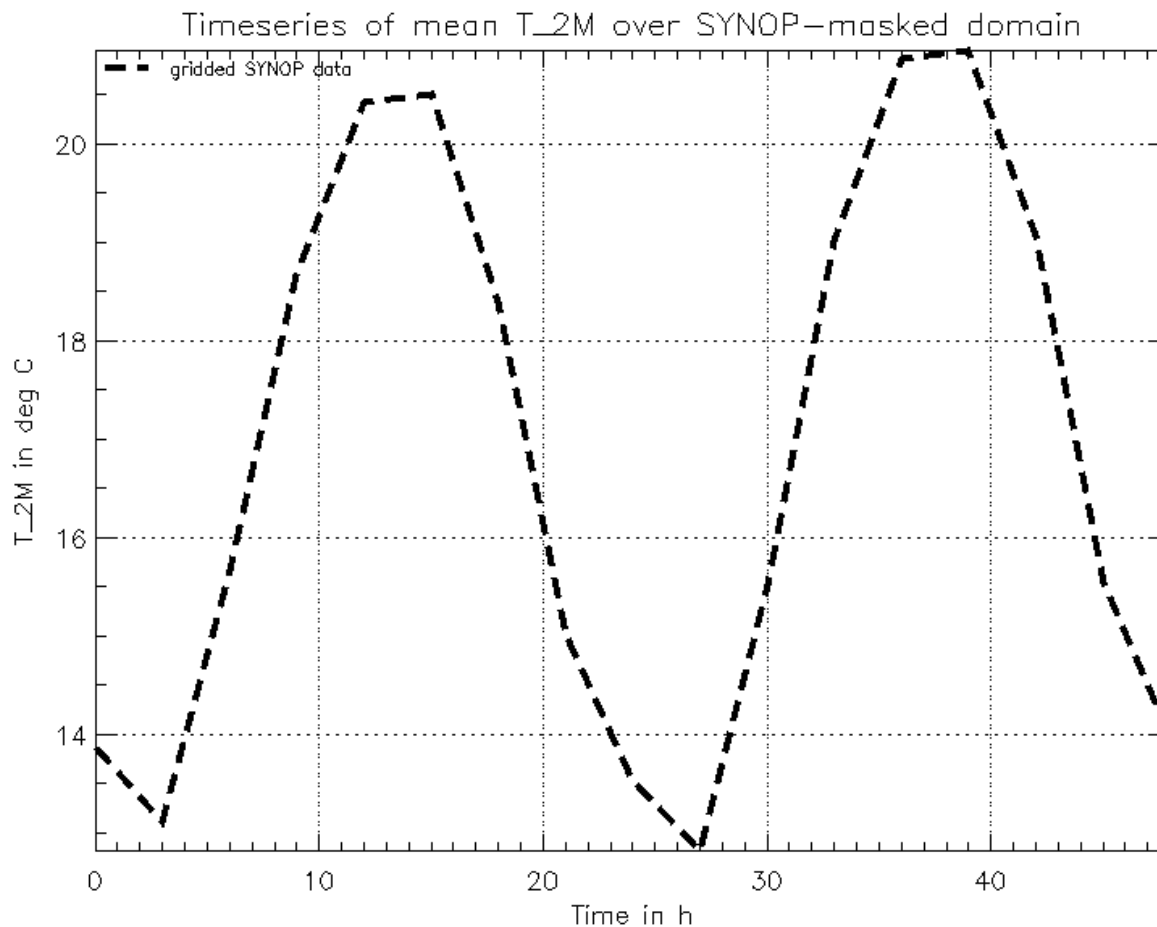
$$\text{width factor} = \frac{\Delta \rho_c}{\bar{\rho}_c} \leq 2$$

3) Include grid scale q_s , q_g into radiative calculations

- Very simplistic methodology:
 - Treat q_s and q_g exactly the same way as q_i , except that for R_e , the maximum value of the validity range of the parameterizations is used, which is $70 \mu\text{m}$.
- This is certainly too simple and has to be improved in the future!
(Careful) extrapolation of the fits, taking into account correct (asymptotic) large-size limit behaviour

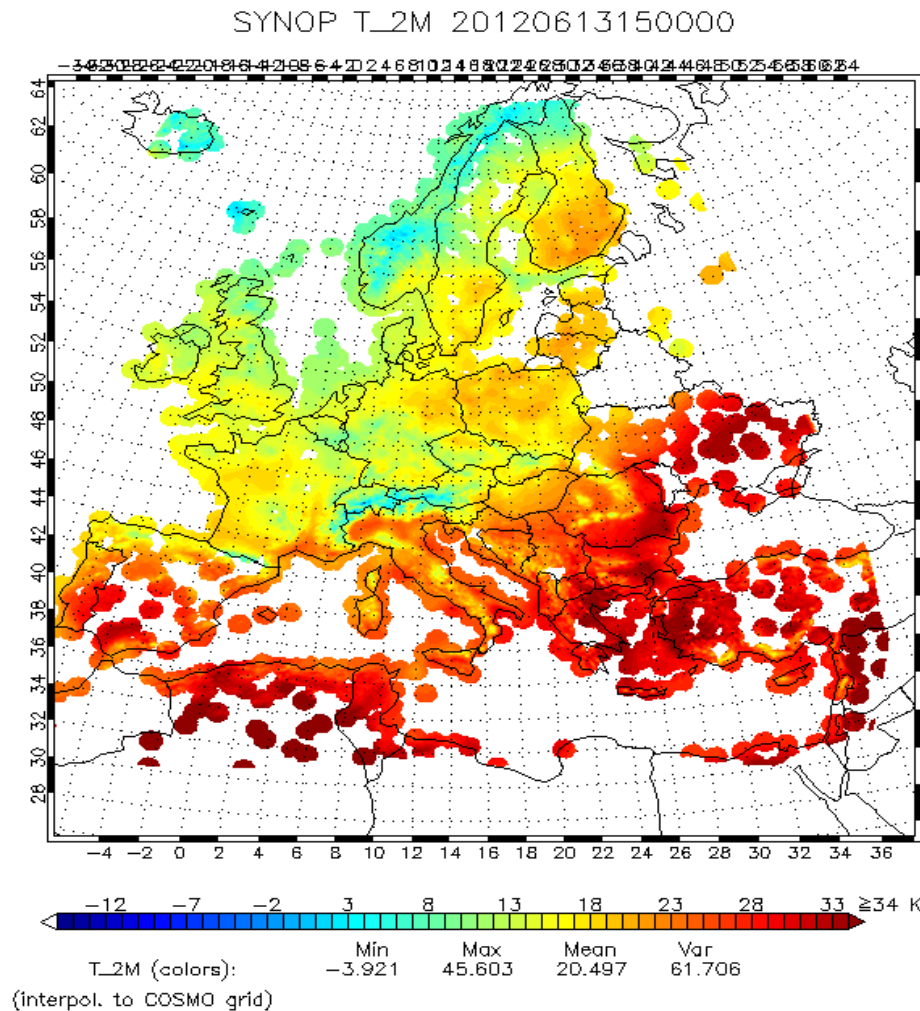
Case study: 13.6.2012 (C-EU)

Cloudy day in central Europe, cf. pictures from the beginning of the talk!



Gridded
synop

Gridded T_{2M} from synop station data (COSMO-EU)



Voronoi-interpolation,
based on triangulation.

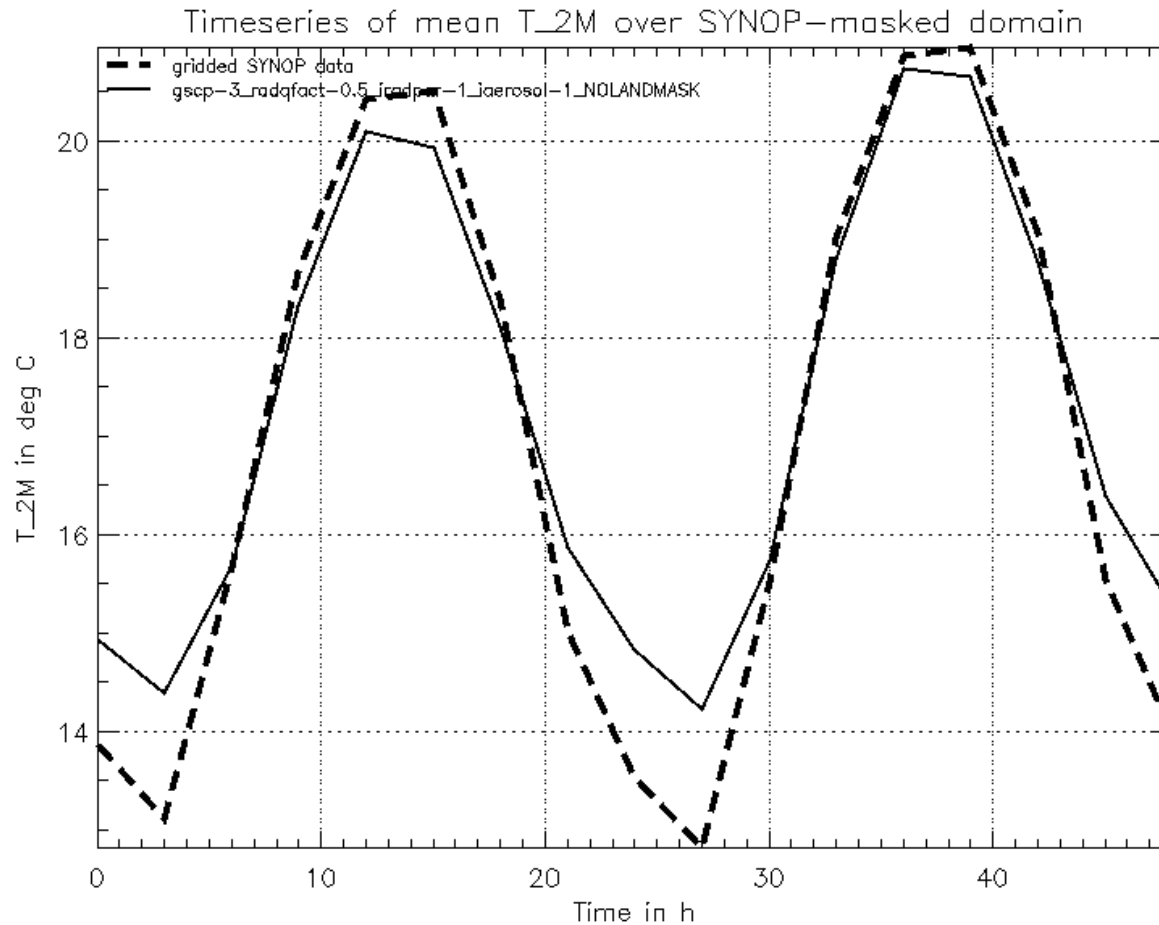
Height correction using
standard atmosphere
gradient.

Max. distance from the
next synop station 70 km.

⇒ Synop data on the model
grid, suitable for
computing, e.g., bias and
rms.

Case study: 13.6.2012 (C-EU)

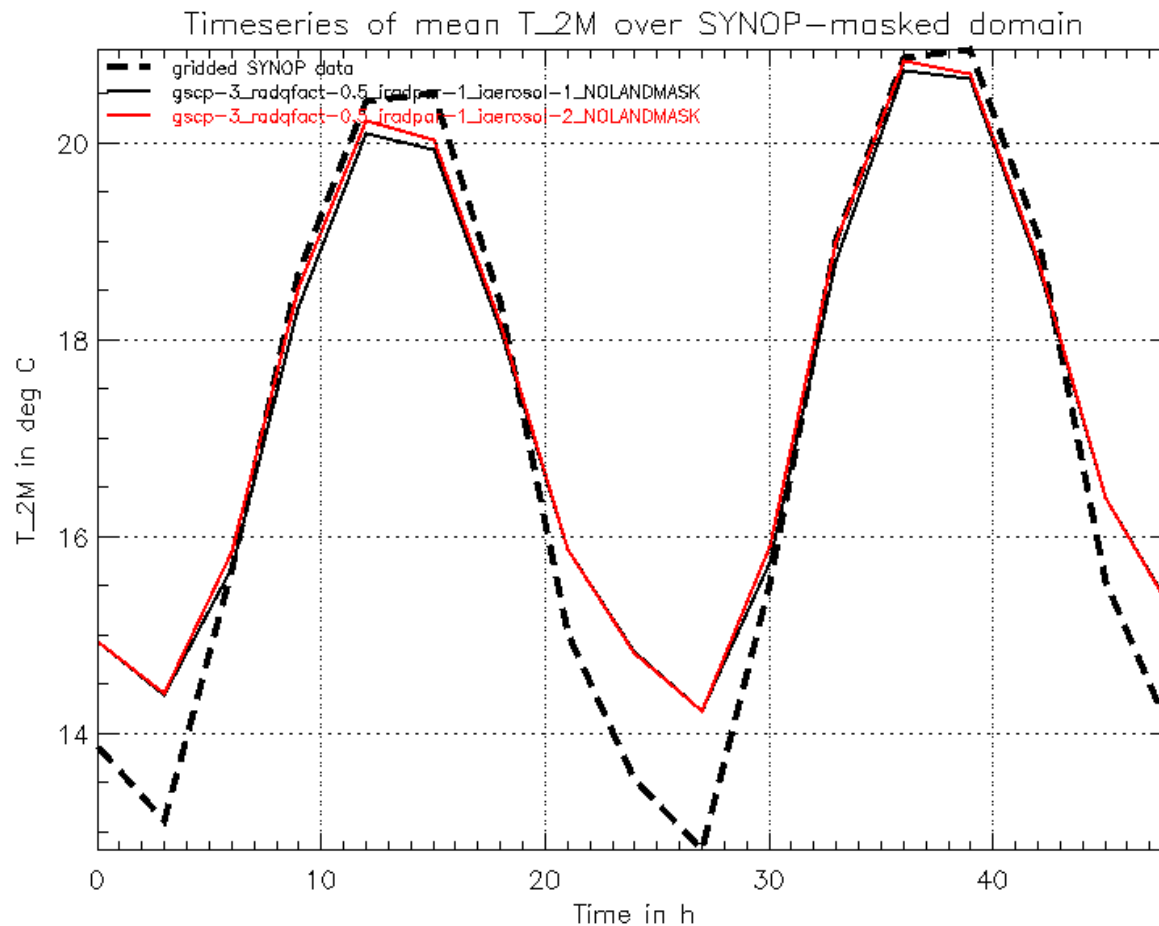
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control

Case study: 13.6.2012 (C-EU)

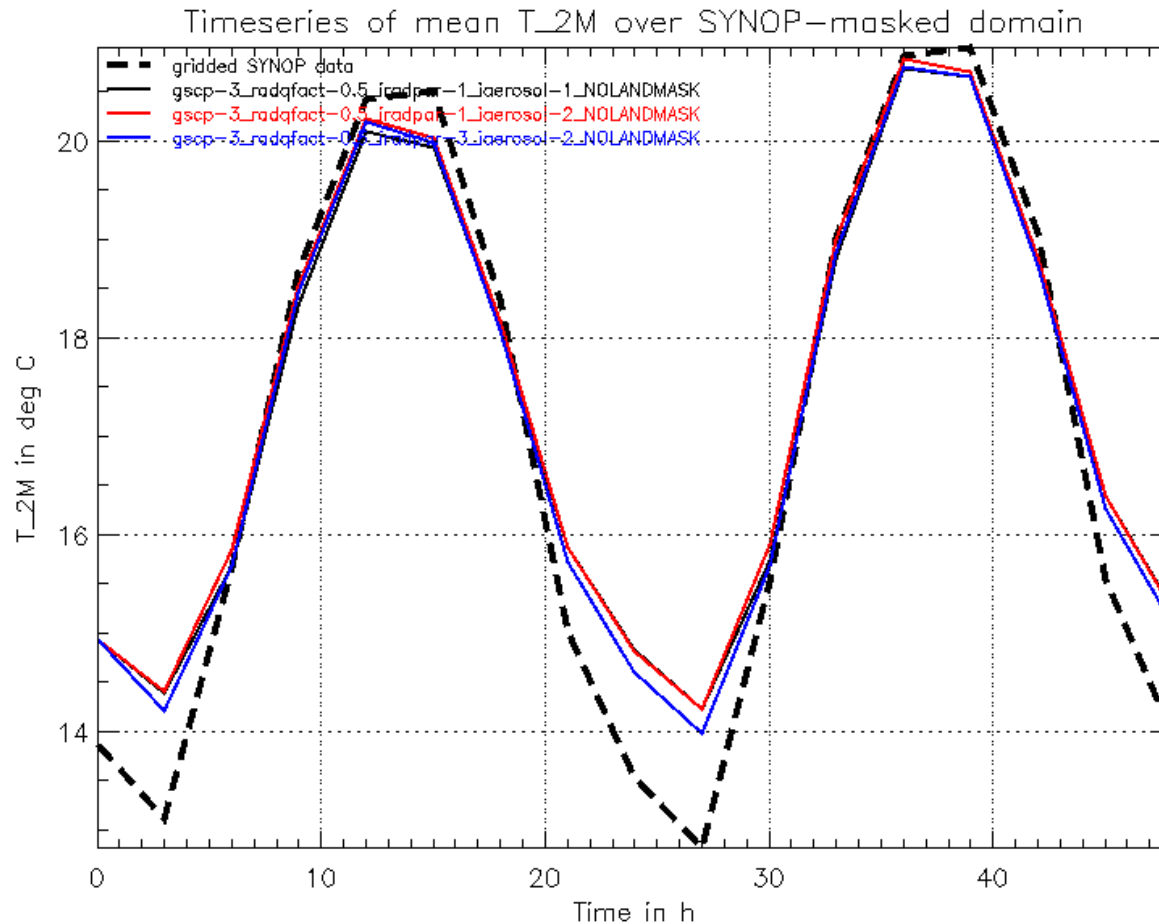
Cloudy day in central Europe, cf. pictures from the beginning of the talk!



+ Tegen
aerosol
climatology
instead of
Tanre

Case study: 13.6.2012 (C-EU)

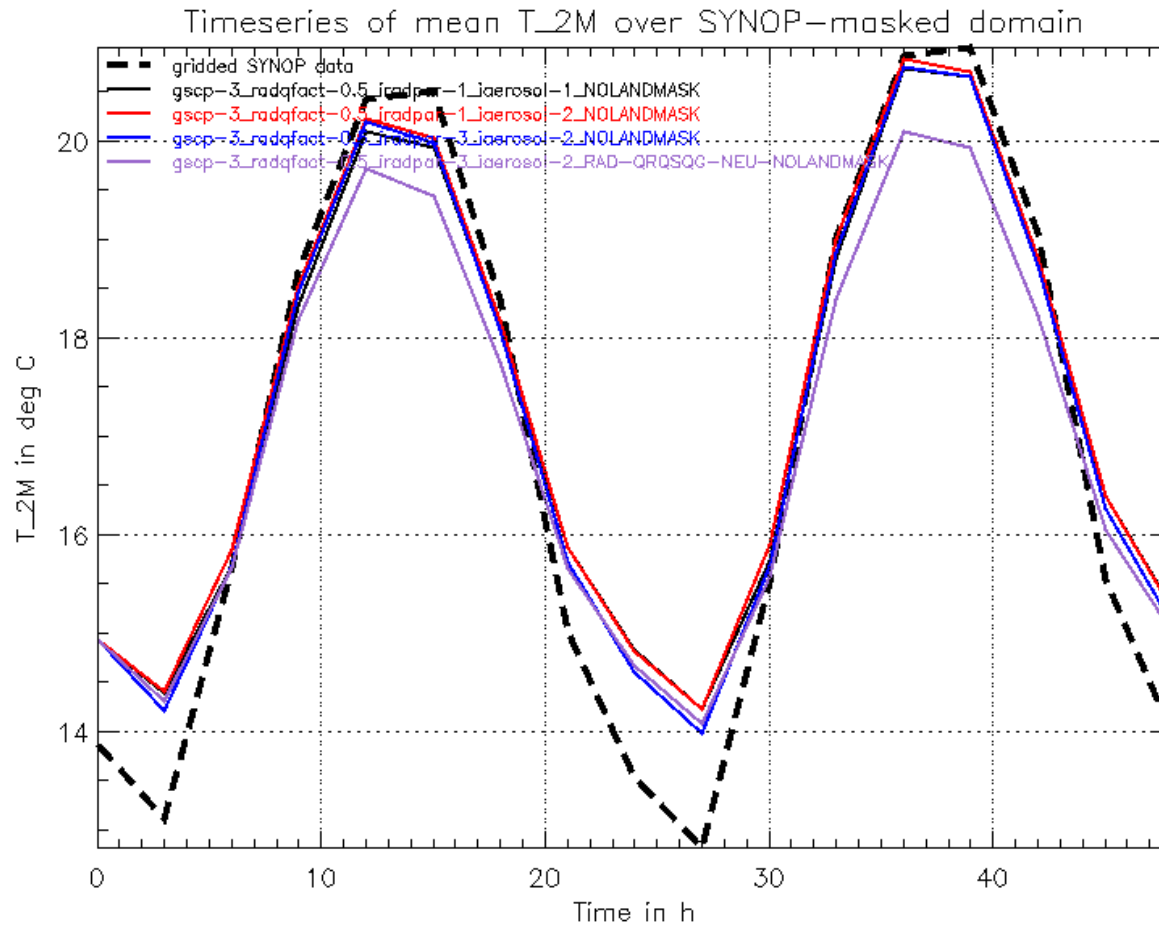
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+ new
optical
properties

Case study: 13.6.2012 (C-EU)

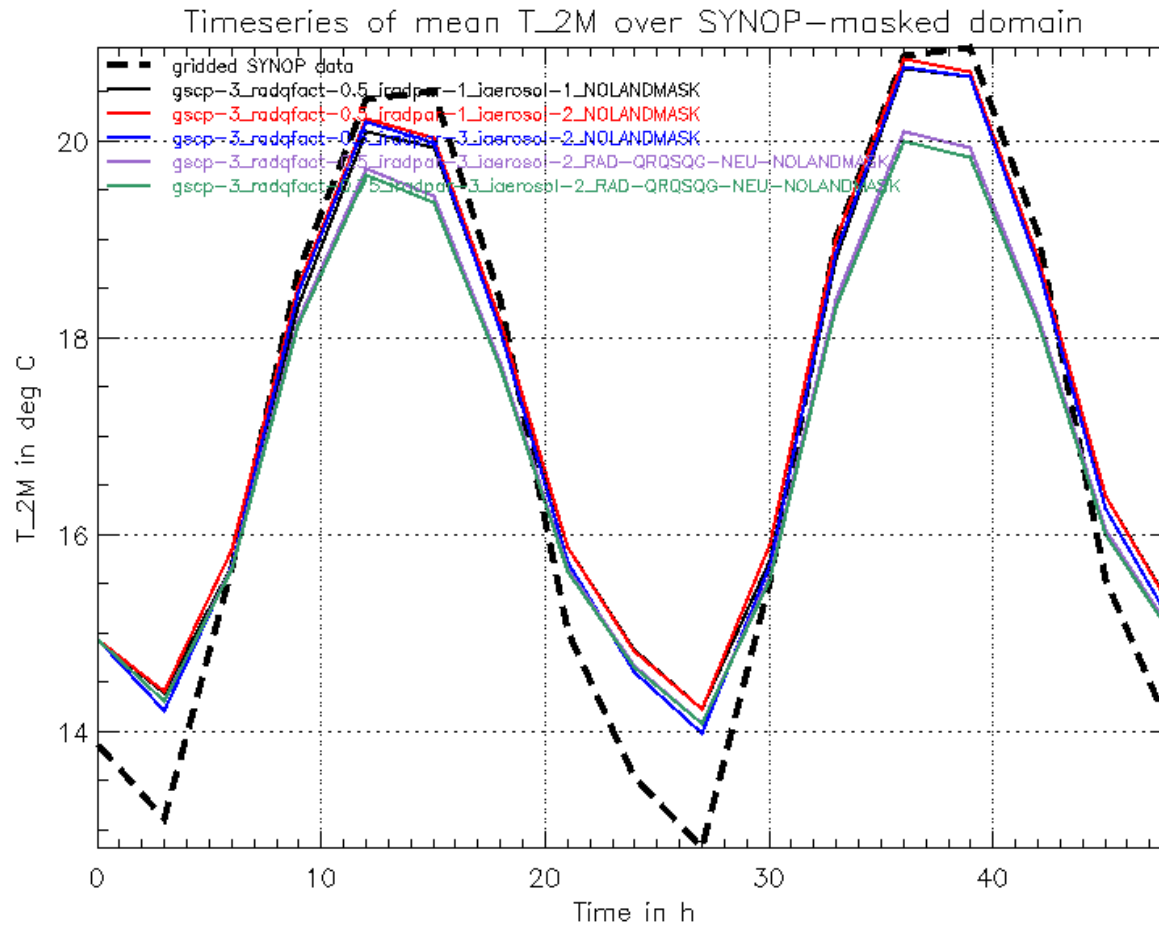
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+ qs, qg

Case study: 13.6.2012 (C-EU)

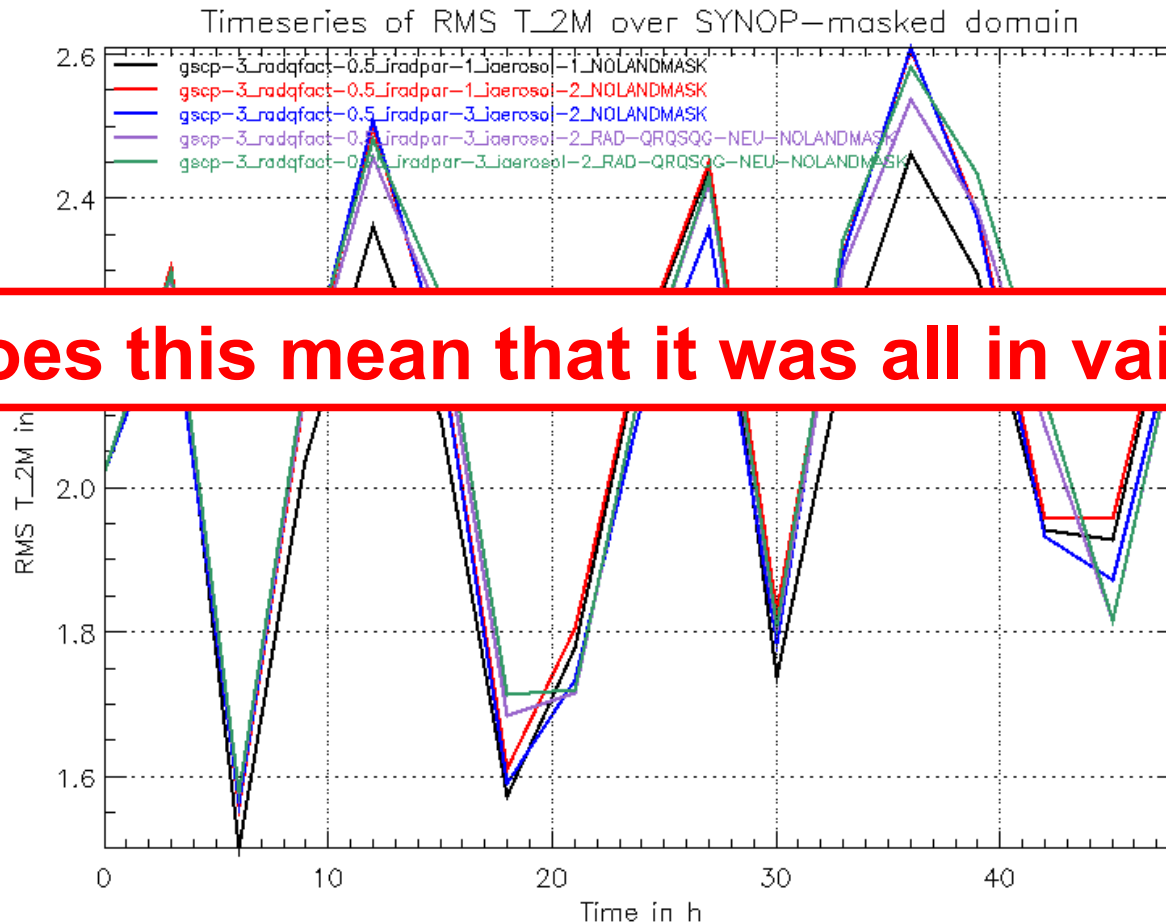
Cloudy day in central Europe, cf. pictures from the beginning of the talk!



$k = 0.75$
instead of 0.5

Case study: 13.6.2012 (C-EU)

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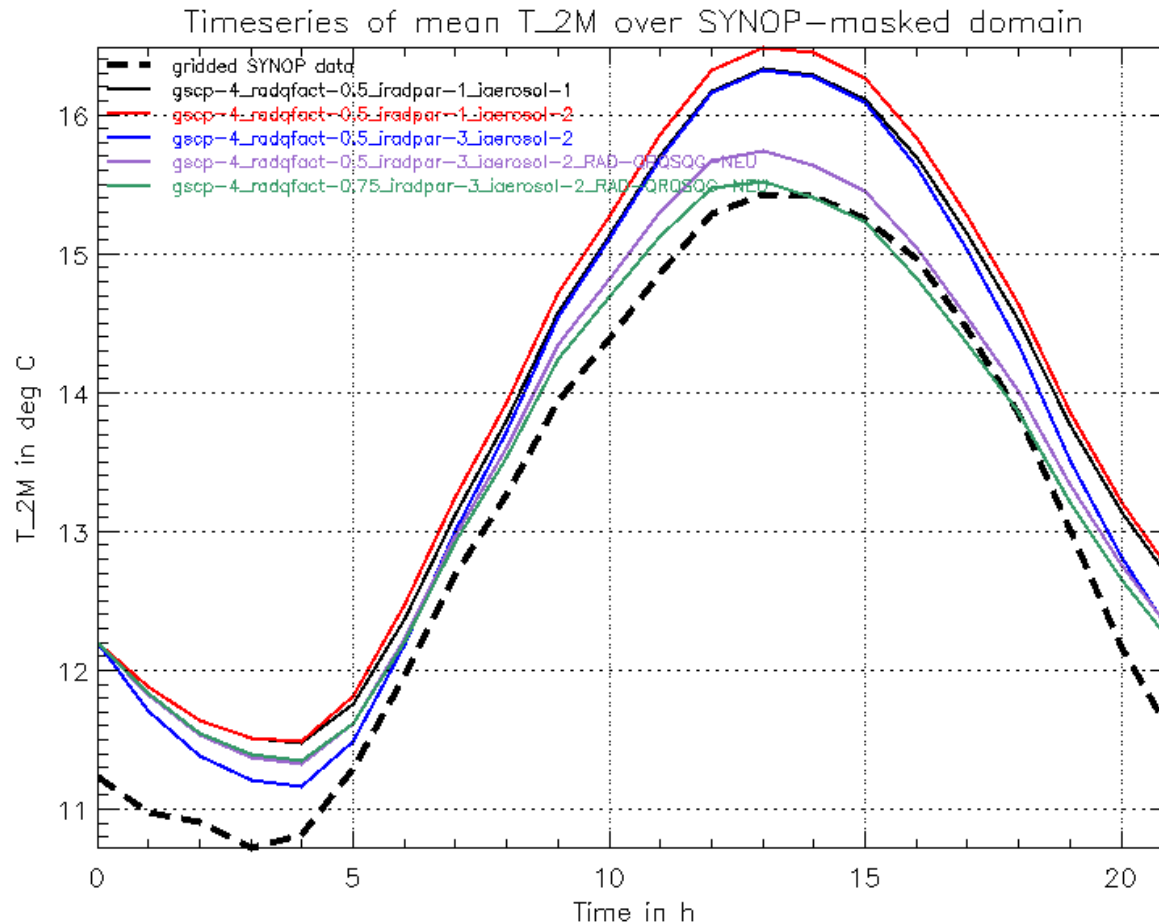
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Does this mean that it was all in vain?

Case study: 13.6.2012 (C-DE)

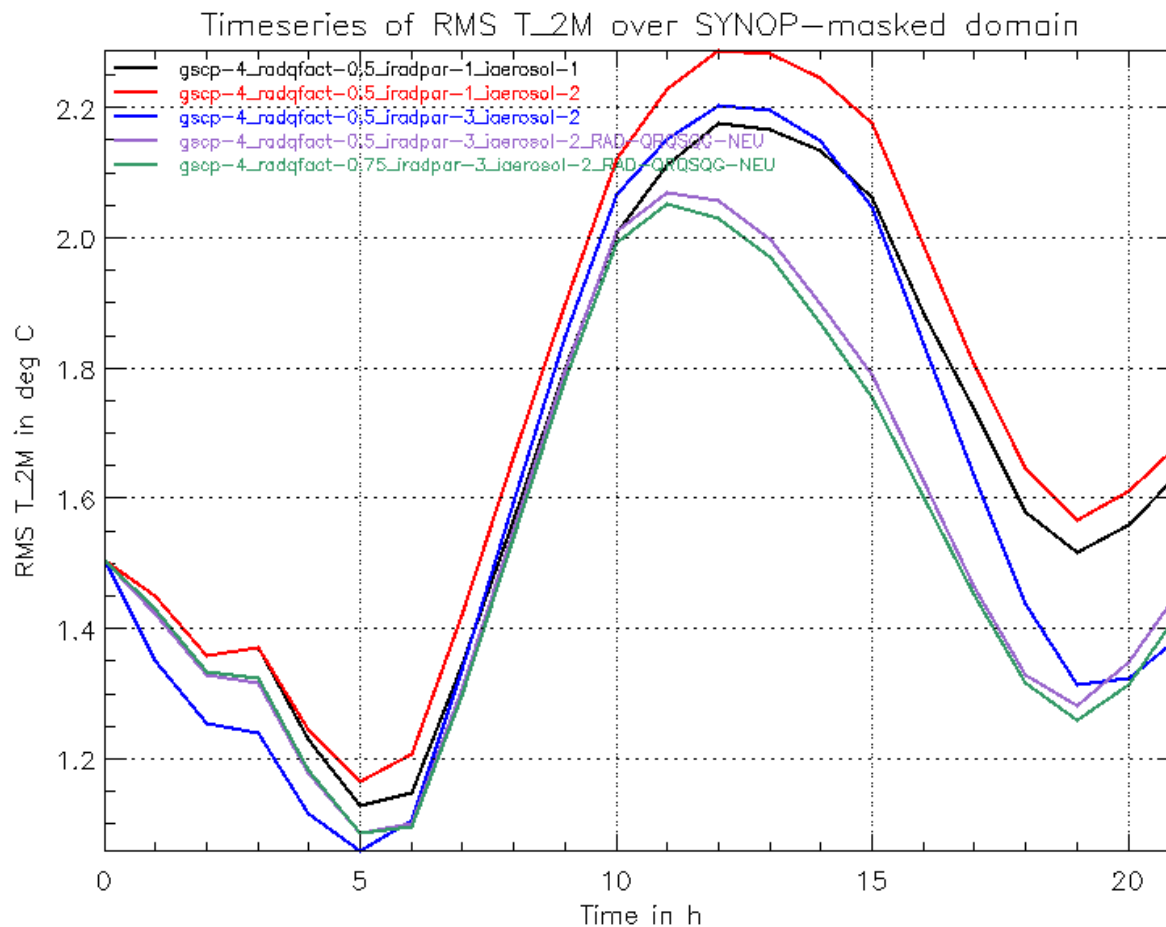
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NO!



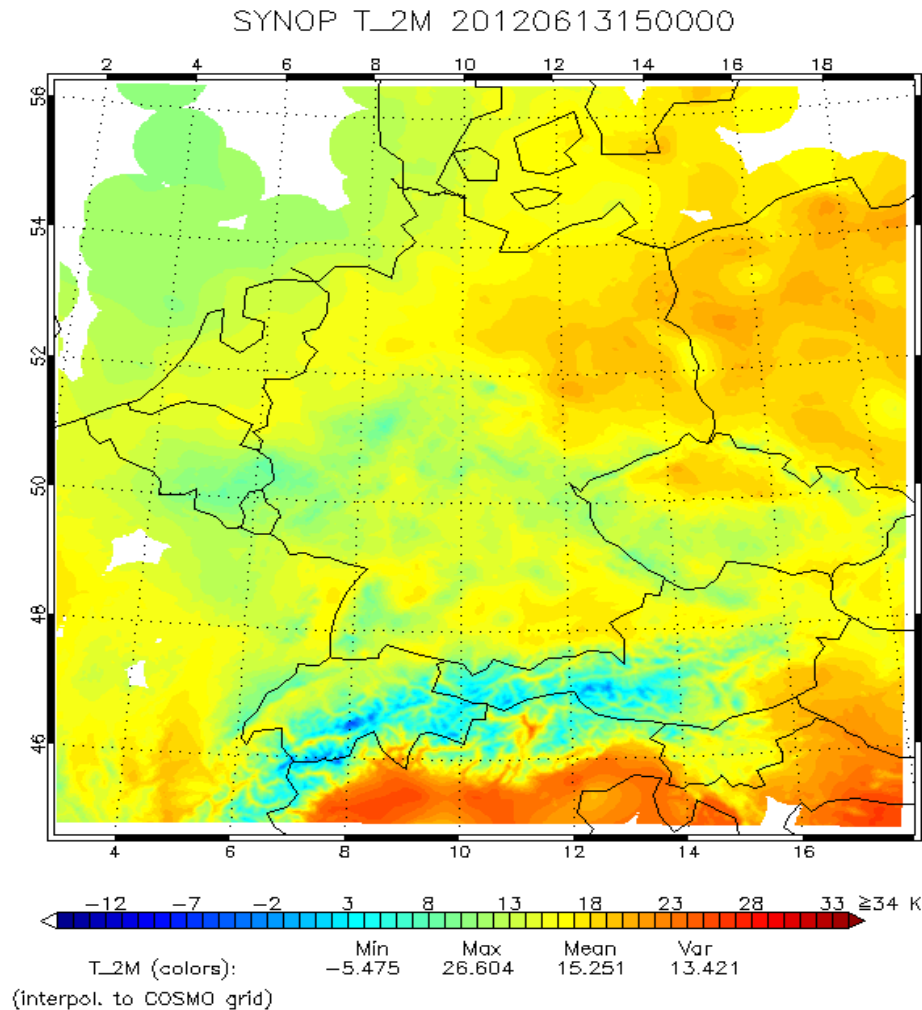
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Cloudy day in central Europe, cf. pictures from the beginning of the talk!



- Changes in cloud radiation coupling can lead to big changes of T_{2M} and possibly other model variables. This gives us a pretty big handle on the model!
- Sensitivities: The implemented R_e -parameterisations make the ice clouds optically thinner in the visible and infrared, therefore increased shortwave heating and longwave cooling in the presence of clouds. Including q_s/q_g and increasing factor k counteract, the clouds get optically thicker at all wavelengths, so T_{max} during day is reduced.
- However, entire model currently tuned to the previous method of cloud radiation coupling (SGS cloud diagnostics, ...). Therefore, to uncover possible beneficial effects of the presented new method will require extensive re-tuning of the model!
- That will be a long process!

Gridded T_{2M} from synop station data (COSMO-DE)



Voronoi-interpolation,
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