

Towards the implementation of a transient gravity wave drag parametrization in ICON

Gergely Bölöni, Ulrich Achatz, Bruno Ribstein

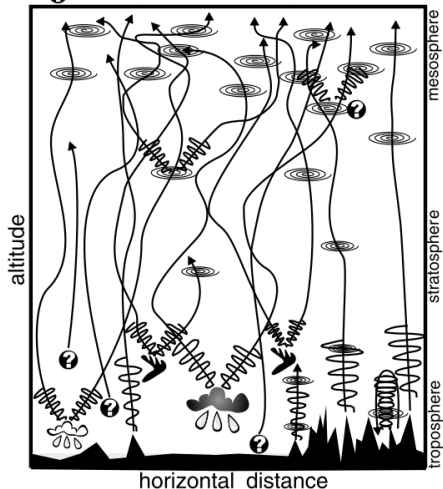
with contributions from:

Sebastian Borchert, Jewgenija Muraschko, Christine Sgoff, Junhong Wei



Motivation

- Gravity Wave Breaking and Drag
- Gravity Wave Group Propagation (Ray) Path
- Gravity Wave Amplitudes and Wave forms
- Jet Stream Instabilities
- Convection/Thunderstorms
- Orography
- Other Unspecified Sources of Gravity Waves



Atmospheric gravity waves (GW)

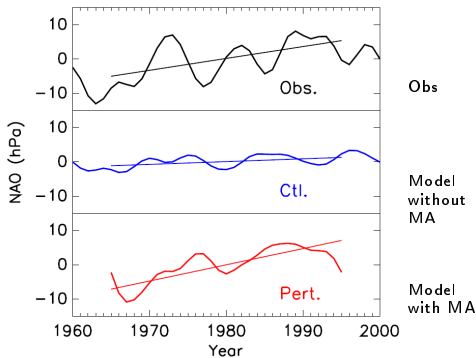
- main sources: orography, convection, jets/fronts
- mainly vertical energy (momentum) transport with $\vec{c}_g \Rightarrow$ interaction with the large scale flow ("drag")
- wave breaking \Rightarrow turbulence, dissipation, energy transfer to large scale flow ("drag")

(Kim et al., 2003)

Motivation

Importance of atmospheric GWs in weather & climate

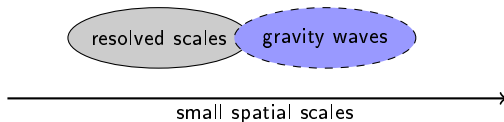
- Mesospheric jet reversal, summer cold pole (*Holton, 1983*)
- Strong effect of GWs on North Atlantic Oscillation (NAO) (*Scaife et al., 2005*)
- Models with middle atmosphere (MA) allow for more accurate GW dynamics \Rightarrow better NAO patterns



Scaife et al. (2005)

Motivation

Parametrization of atmospheric GWs



- GWs are not fully resolved by GCMs and NWP models \Rightarrow parametrization \Rightarrow (Wentzel–Kramers–Brillouin) WKB theory
- Currently used parametrizations: steady state approximation \Rightarrow instantaneous propagation through constant resolved flow \Rightarrow instantaneous drag via wave breaking only!
- Proposal for improvement: weakly-nonlinear coupling between the GW and the resolved flow \Leftrightarrow transient propagation \Leftrightarrow continuous drag on the resolved flow during propagation + drag through wave breaking

WKB theory

Wave resolving system (2-D Euler equations, no rotation):

$$\frac{Du}{Dt} + c_p \theta \frac{\partial \pi}{\partial x} = 0$$

$$\frac{Dw}{Dt} + c_p \theta \frac{\partial \pi}{\partial z} + g = 0$$

$$\frac{D\theta}{Dt} = 0$$

$$\frac{D\pi}{Dt} + \frac{\kappa}{1-\kappa} \pi \left(\frac{\partial u}{\partial x} + \frac{\partial w}{\partial z} \right) = 0$$

with $\frac{D}{Dt} = \frac{\partial}{\partial t} + u \frac{\partial}{\partial x} + w \frac{\partial}{\partial z}$

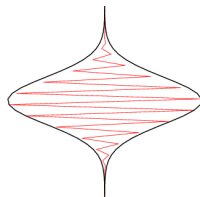
Exner pressure $\pi = (p/p_0)^\kappa$

Pot. temperature $\theta = T(p_0/p)^\kappa = T/\pi$

$$\kappa = R/c_p$$

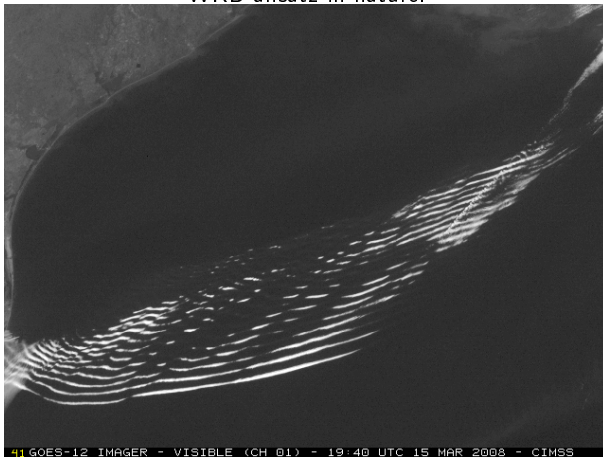
Simplification ingredients:

- Decomposition of the fields: $f = f_b + f_w$
- WKB ansatz: $f_w(x, z, t) = \text{Re} F_w(Z, T) e^{i[kx + \frac{\phi(Z, T)}{\epsilon}]}$
with $Z = \epsilon z, T = \epsilon t, m = \partial\phi/\partial Z$ and $\omega = -\partial\phi/\partial T$
- Scaling for the gravity waves: $\epsilon = \lambda_{wave}/\lambda_{background} \ll 1$



WKB theory

WKB ansatz in nature:



(cimss.ssec.wisc.edu)

WKB theory

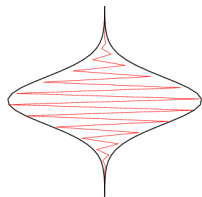
Wave resolving equations (2-D Euler equations, no rotation):

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WKB theory

Order analysis in $\epsilon \Rightarrow$ GW parametrization

Wave field

Mean flow

WKB theory: transient coupled system

$$\begin{aligned}\frac{d_g z}{dt} &= \mp \frac{Nkm}{(k^2 + m^2)^{3/2}} \equiv c_{gz} & \frac{\partial u_b}{\partial t} &= -\frac{1}{\bar{\rho}} \frac{\partial}{\partial z} (kc_{gz} \mathcal{A}) \\ \frac{d_g m}{dt} &= \mp \frac{k}{(k^2 + m^2)^{1/2}} \frac{dN}{dz} - k \frac{d u_b}{dz} \equiv \dot{m} \\ \frac{d_g \mathcal{A}}{dt} &= -\mathcal{A} \frac{\partial c_{gz}}{\partial z} \quad \left(\frac{d_g}{dt} = \frac{\partial}{\partial t} + c_{gz} \frac{\partial}{\partial z} \right)\end{aligned}$$

Steady state WKB theory: instantaneous decoupled system

$$\begin{aligned}\frac{d_g z}{dt} &= \mp \frac{Nkm}{(k^2 + m^2)^{3/2}} \equiv c_{gz} & \frac{\partial u_b}{\partial t} &= -\frac{1}{\bar{\rho}} \frac{\partial}{\partial z} (kc_{gz} \mathcal{A}) \\ \frac{d_g m}{dt} &= 0 \\ \frac{\partial \mathcal{A}}{\partial t} &= 0 \iff c_{gz}(z) \mathcal{A}(z) = \text{const.}\end{aligned}$$

\Rightarrow no wave-mean-flow interaction! \Rightarrow
Drag sources: ONLY! wave breaking (constraining $\mathcal{A}(z)$)

WKB theory

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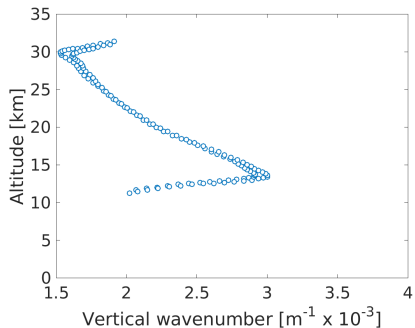
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Numerical implementation: toymodel

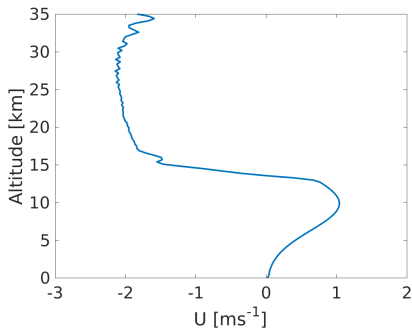
GW field

Ray positions in phase space



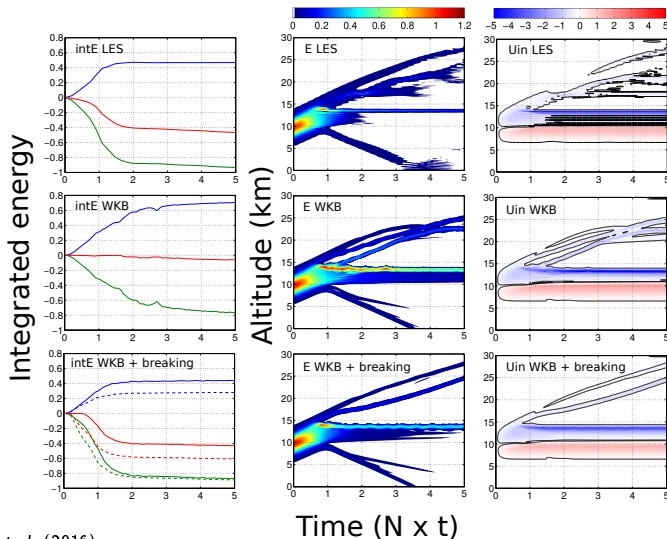
Meanflow

Induced large scale wind



Numerical implementation: toymodel

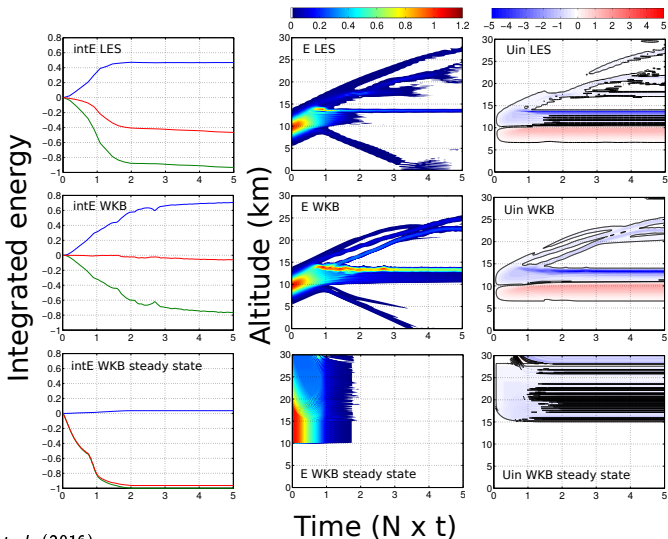
Static instability ($\lambda_x = \lambda_z = 1km$)



Böläni et al. (2016)

Numerical implementation: toymodel

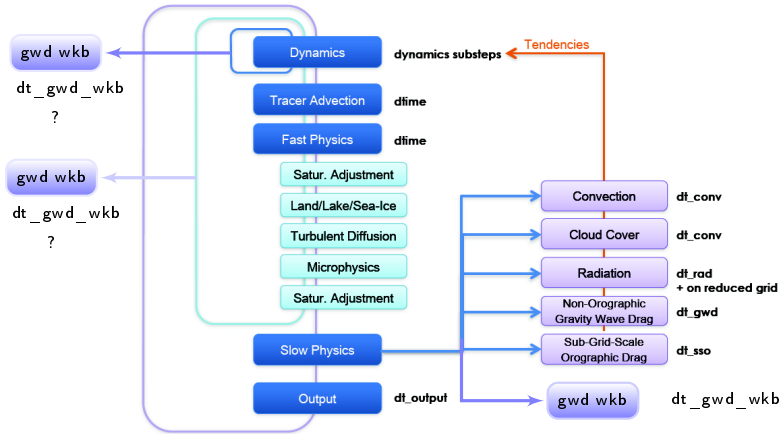
Static instability ($\lambda_x = \lambda_z = 1km$)



Böläni et al. (2016)

Numerical implementation: ICON

Concept

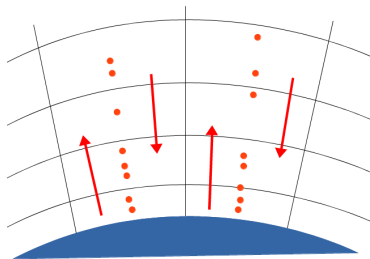


(Original courtesy: DWD, ICON Training 2015)

Numerical implementation: ICON

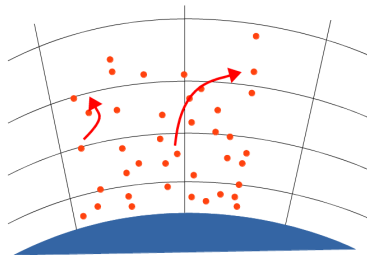
Concept

1D framework



Fits well to the current MPI communicator

3D framework



Requires new MPI communication style for Lagrangian particles \Rightarrow later...

Numerical implementation: ICON

Primary code validation

WKB in ICON

WKB in toymodel

- low resolution: $dx \approx 1250km$ (R02B01), $dz \approx 200m$
- no orography, simple stratification, steady state initial meanflow
- dynamical core off (`ldynamics=.F.`), no rotation (`lcoriolis=.F.`)
- all other physical parametrizations switched off
- known cases, uniform source over the Globe \Rightarrow "single column"

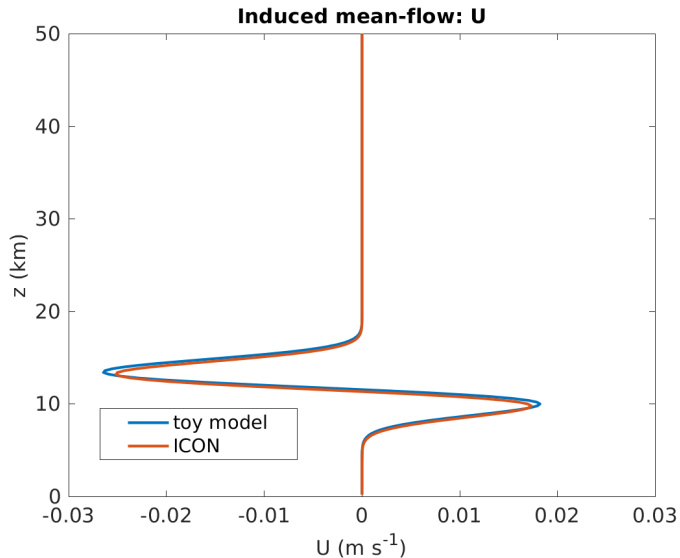
Moving towards realism...

WKB in ICON

OPER ICON / OBS

- $dx \approx 160km$ (R02B04), $dx \approx 80km$ (R02B05), $dz \approx 50 - 350m$
- switch on dynamical core (`ldynamics=.T.`) and rotation (`lcoriolis=.T.`)
 \Rightarrow interaction of GWs and inertial waves
- real orography, same GW sources as in OPER ICON NWP (*Orr et al., 2010*)
- comparison of zonal means, etc.

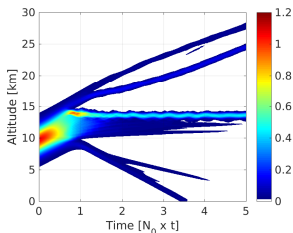
Numerical implementation: ICON



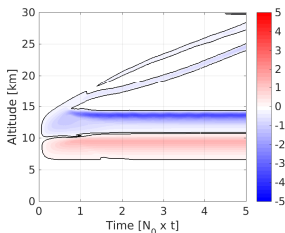
Numerical implementation: ICON

Static instability ($\lambda_x = \lambda_z = 1km$)

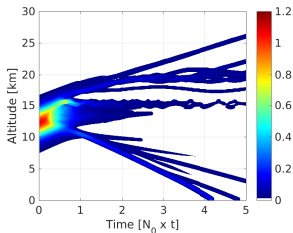
Energy (toymodel)



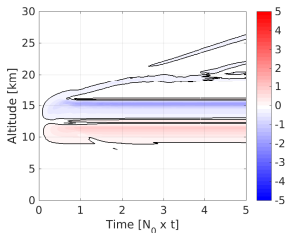
Induced U (toymodel)



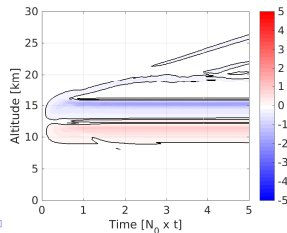
Energy (ICON)



Induced U (ICON)



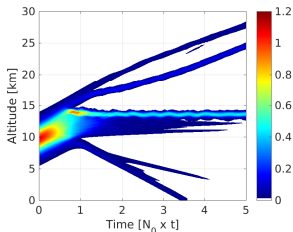
Induced V (ICON)



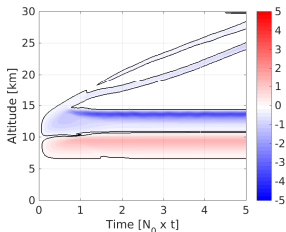
Numerical implementation: ICON

Static instability ($\lambda_x = \lambda_z = 1km$)

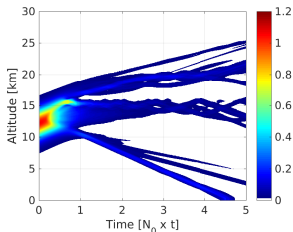
Energy (toymodel)



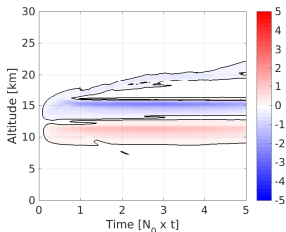
Induced U (toymodel)



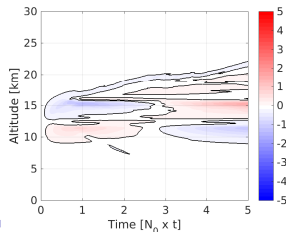
Energy (ICON rotation)



Induced U (ICON rotation)

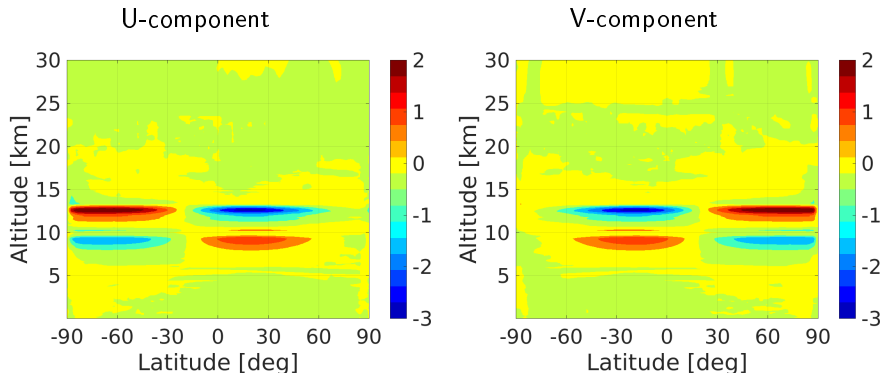


Induced V (ICON rotation)



Numerical implementation: ICON

Static instability ($\lambda_x = \lambda_z = 1km$)



Conclusions

- A **new GW drag parametrization** proposed (**MS-GWAM¹**): *full* WKB theory \Leftrightarrow *steady state* WKB theory.
- Based on idealized MS-GWAM simulations **transient wave-meanflow interactions are more important than wave breaking**. Current GW drag schemes are based only on the latter process.
- An **MS-GWAM 1.0** is **implemented** and technically validated in **ICON (NWP)**. It is to be gradually extended: more realistic sources, 1D \rightarrow 3D, etc.

¹Multi Scale Gravity Wave Model

References

Achatz, U., R. Klein, F. Senf (2010), Gravity waves, scale asymptotics, and the pseudo-incompressible equations. *J. Fluid Mech.*, 141(663), 120–147, DOI:10.1017/S0022112010003411

Böläni, G., Ribstein, B., Muraschko, J., Sgoff, C., Wei, J. and Achatz, U., 2016: The Interaction between Atmospheric Gravity Waves and Large-Scale Flows: An Efficient Description beyond the Nonacceleration Paradigm. *J. Atmos. Sci.*, 73, 4832 - 4852, DOI:10.1175/JAS-D-16-0069.1.

Holton, J. R. (1983), The influence of gravity wave breaking on the general circulation of the middle atmosphere, *J. Atmos. Sci.*, 40, 2497–2507.

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Scaife, A. A., J. R. Knight, G. K. Vallis, and C. K. Folland (2005), A stratospheric influence on the winter NAO and North Atlantic surface climate, *Geophys. Res. Lett.*, 32, L18715, DOI:10.1029/2005GL023226.