



An upper-atmosphere extension of the ICON-model to study gravity waves from the troposphere to the lower thermosphere

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Outline

- DFG research group MS-GWaves
 - Our sub-project GWING
- Upper-atmosphere extension of ICON
 - Dynamics
 - Physics
- Plans for 2nd phase of GWING





The DFG research group: *MultiScale dynamics of Gravity Waves** ...

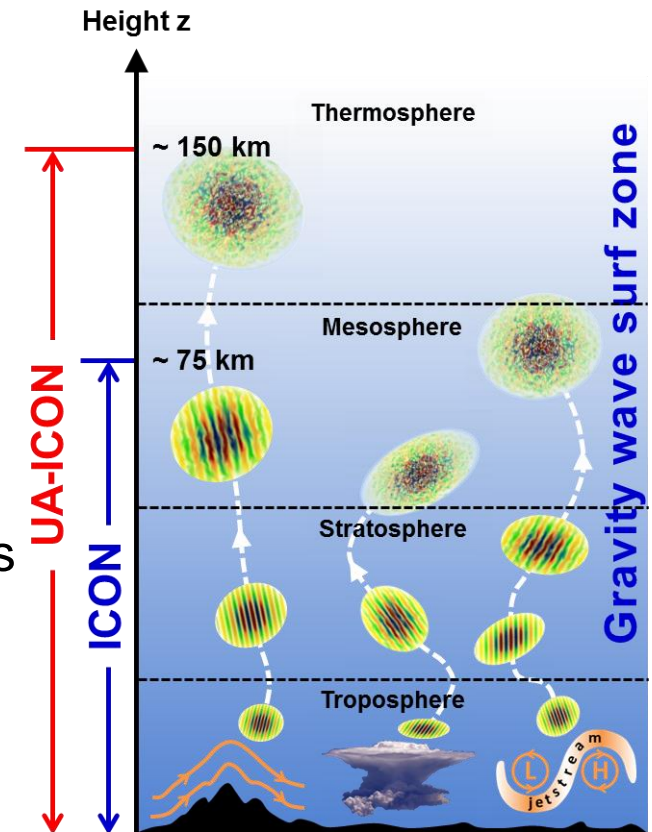
Subject and goals:

- Spatial, temporal and spectral distribution of gravity waves (GWs) in the atmosphere
- Processes causing and controlling the corresponding GW dynamics
- Role of GWs for global atmospheric circulation and its variability
- Improvement of parameterizations of GWs as a sub-grid scale phenomenon in weather prediction and climate models
- Now in its 2nd phase



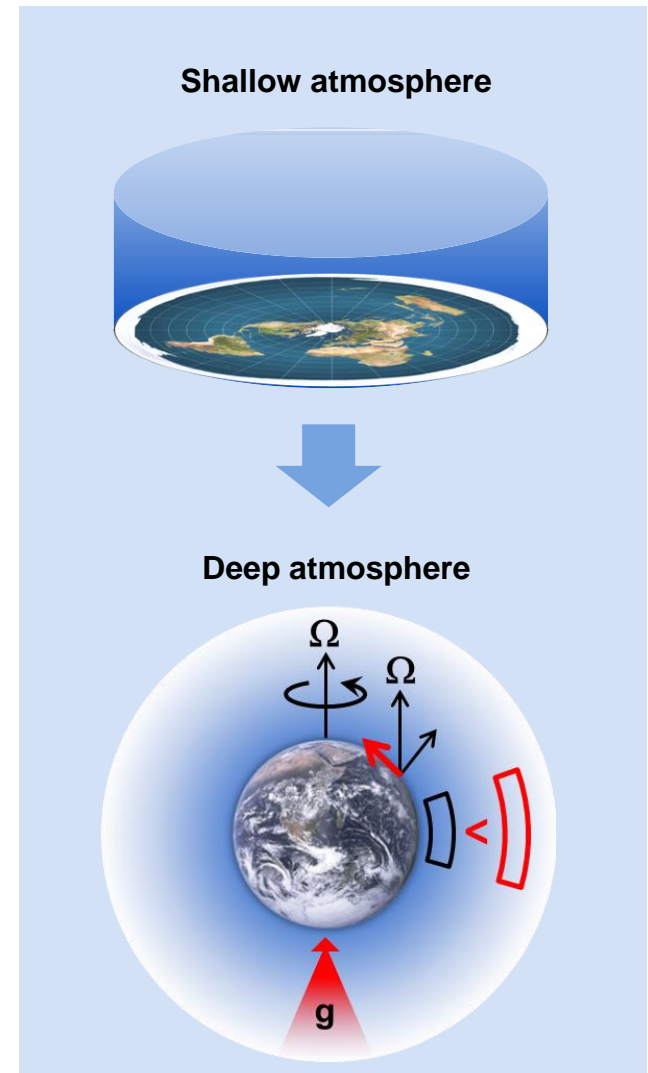
... and our sub-project *Gravity Wave Interactions in the Global atmosphere*

- **Upper-atmosphere configuration: UA-ICON**
 - UA physics package (MPI-M)
 - Dynamics modification (DWD)
 - Evaluation with benchmarks: HAMMONIA model, measurements, ...
- **In collaboration with our partners**
 - Climate simulations and simulations of episodes from the measurement campaigns
 - Investigation and implementation of newly developed GW parameterizations:



Modification of the dynamical core of ICON

- **Shallow-atmosphere approximation:**
 - Standard configuration in ICON
 - Geometrically atmosphere is 'plane'
- **Deep-atmosphere modifications:**
 - Increase of grid cell extension with height
 - Consider Coriolis acceleration due to Ω_h
 - Gravitational acceleration g decreases with height



Testing of modifications (example)

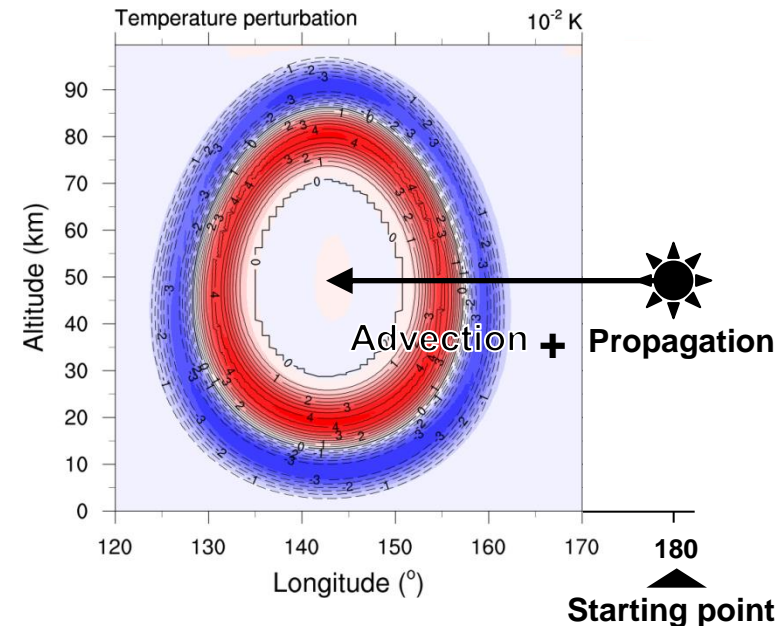
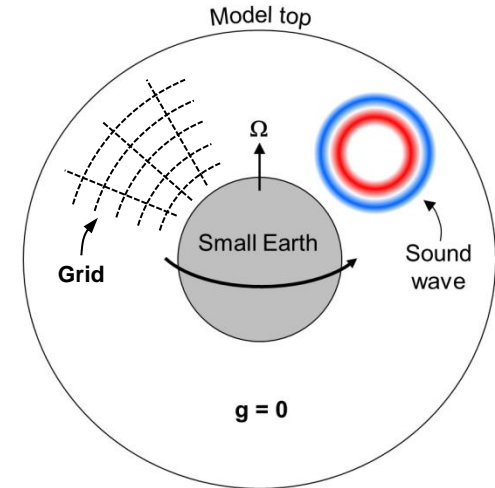
■ Sound wave test case*:

- Homog. atmosphere at rest in abs. frame + spherical sound wave perturbation
- Rotating, small Earth: increase deep-atmosphere effects
- Gravitational acceleration: **OFF**
- Centrifugal acceleration: **ON**

■ Features:

- Analytical solution available
- Test dycore performance on spherically curved grid
- Test conservation of 'balance':

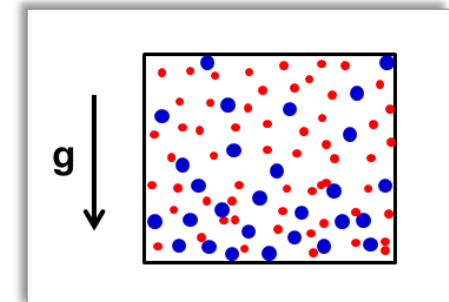
$$dv/dt + \text{Coriolis} + \text{Centrifugal} = 0$$



Implementation of UA-specific physics packages in ICON (at MPI-M)

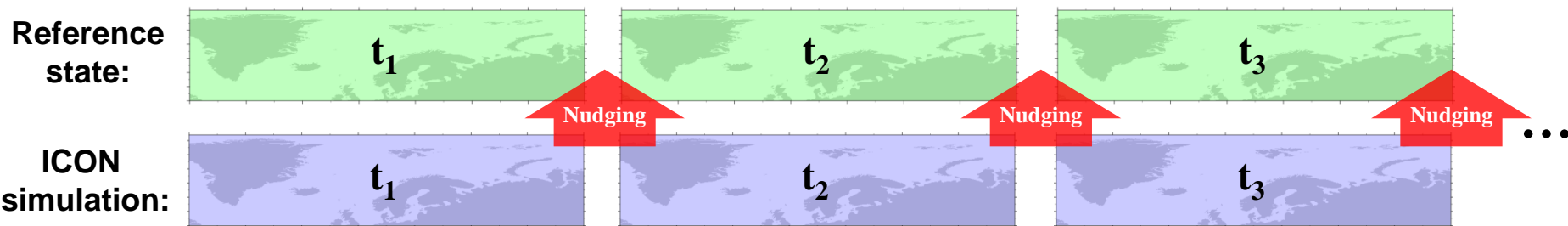
- Physics packages come mainly from the HAMMONIA model (MPI-M, hydrostatic)
- Implementation in (non-hydrostatic) ICON required modification and retuning

<i>Kinetics</i>	
Molecular diffusion	Huang et al. 1998; Banks and Kockarts 1973
Frictional heating	Gill 1982
Gravity wave-induced turbulent diffusion	Hines 1997a,b; Akmaev et al. 1997
<i>Radiation</i>	
Schumann-Runge bands and continuum (O ₂)	Strobel 1978
Extreme ultraviolet (N ₂ , O, O ₂)	Richards et al. 1994
Non-LTE infrared cooling (CO ₂ , NO, O ₃)	Fomichev and Blanchet 1995; Fomichev et al. 1998
NO infrared cooling at 5.3μm	Kockarts 1980
<i>Ionization</i>	
Ion drag	Hong and Lindzen 1976
Joule heating	Hong and Lindzen 1976
<i>Others</i>	
Chemical heating	Climatology from HAMMONIA



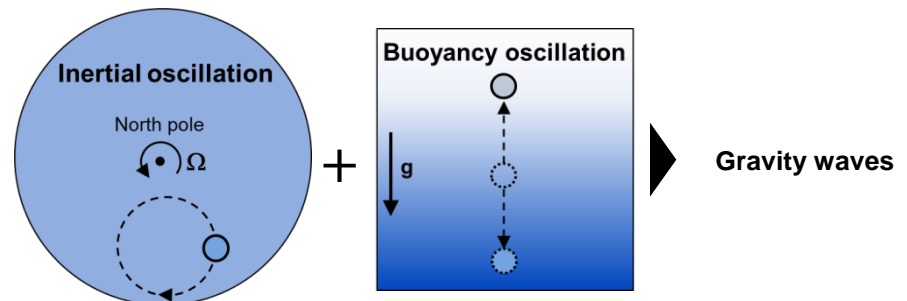
Plans: a nudging option for ICON

- **Motivation:** allow for simulations of contiguous episodes longer than a few days without explicit coupling to a data assimilation scheme, e.g., for comparison with field campaigns

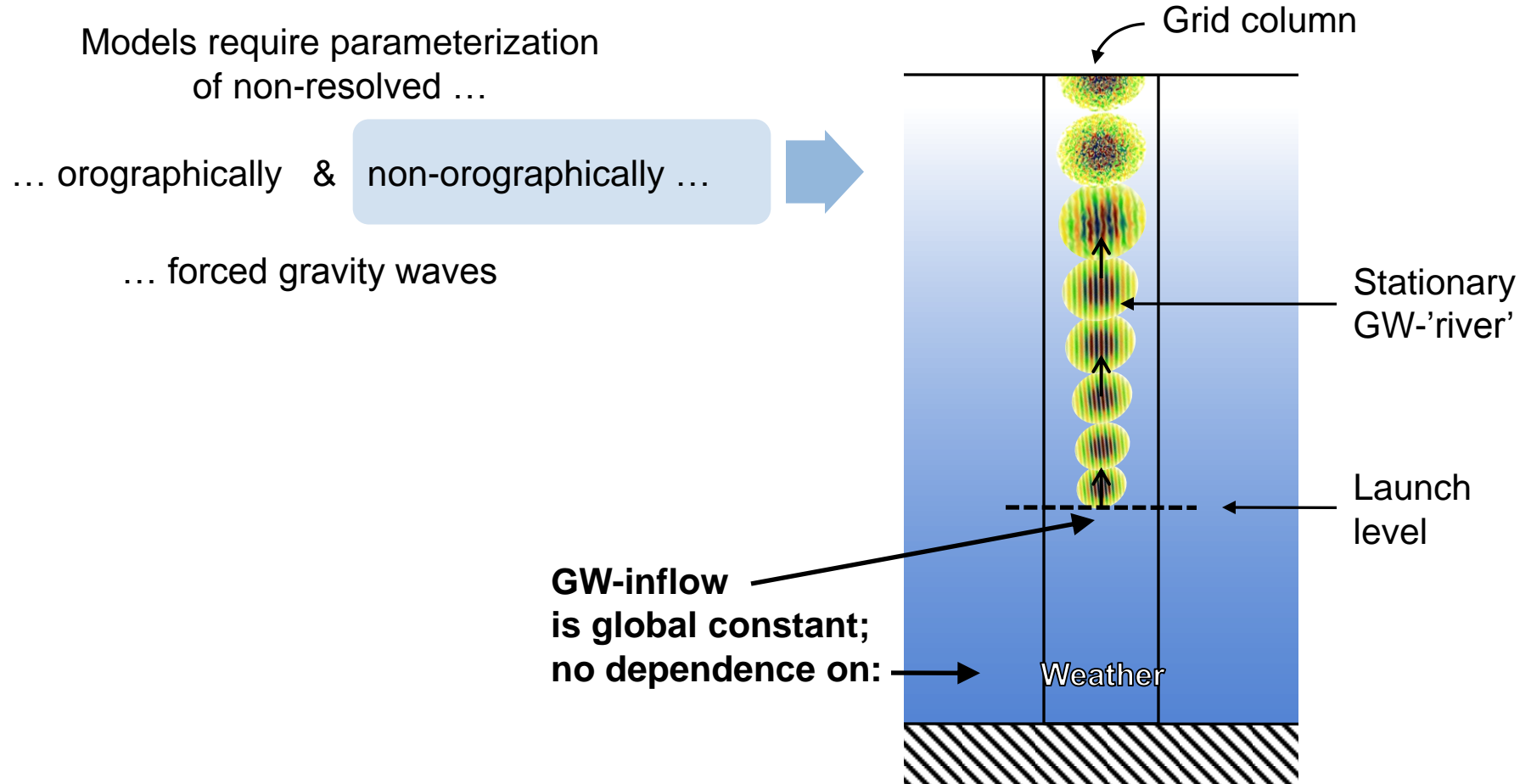


- Preparation & modification of code infrastructure, e.g., for reference data prefetching & reading
- Theoretical considerations: reduce excitation of spurious gravity waves from forcing towards reference state

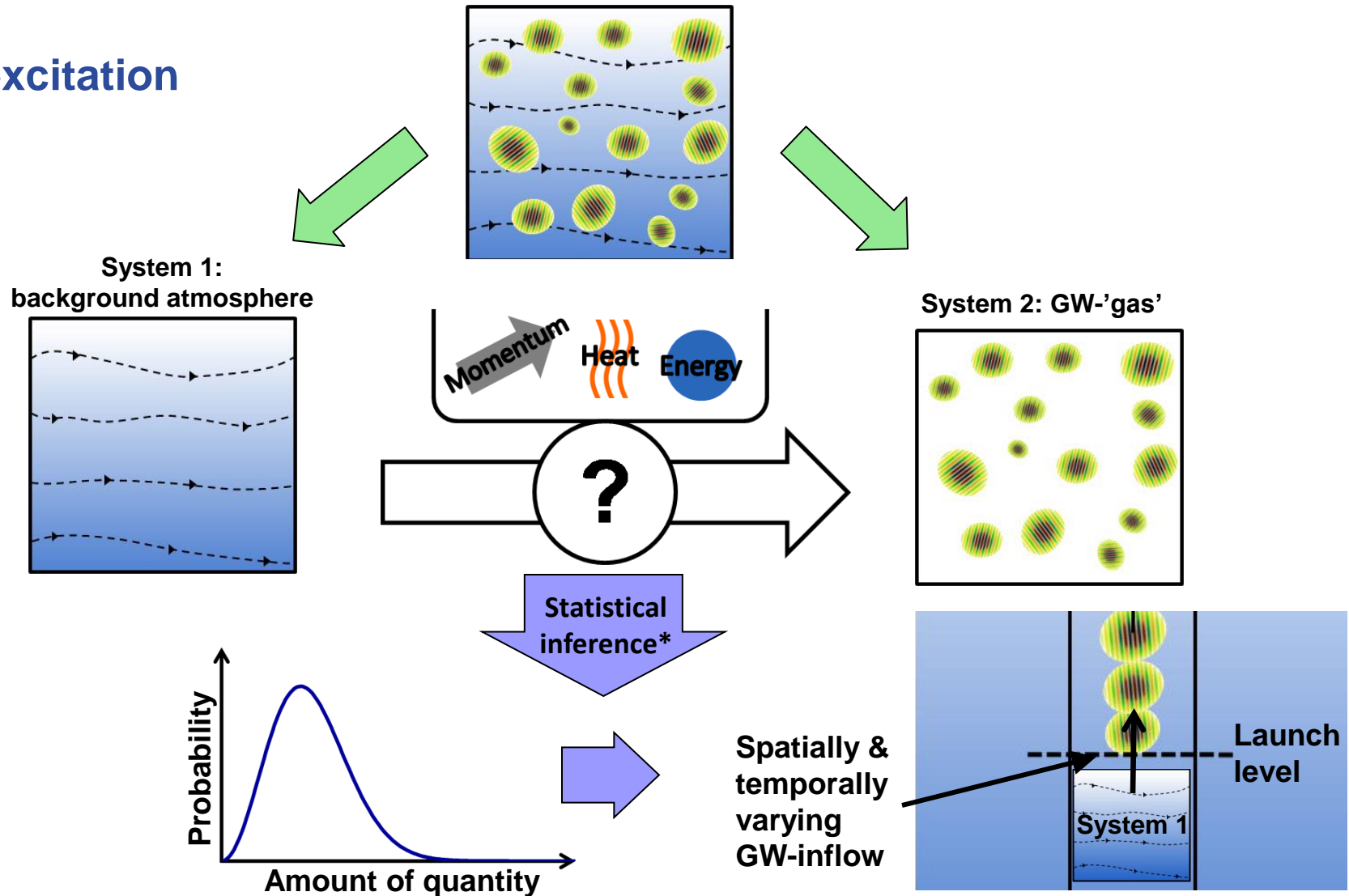
Do not kick off too strong:



Plans: investigate potential of gravity wave parameterizations



GW excitation





Summary

Within the GW research group MS-GWaves:

- Extension of ICON by an upper-atmosphere configuration
 - Deep-atmosphere dynamics
 - Implementation of additional parameterizations for upper-atmosphere physics
 - Evaluation
- Examples for plans in 2nd phase:
Nudging & Potential for improvement of 'standard' GW parameterizations

Thank you!

