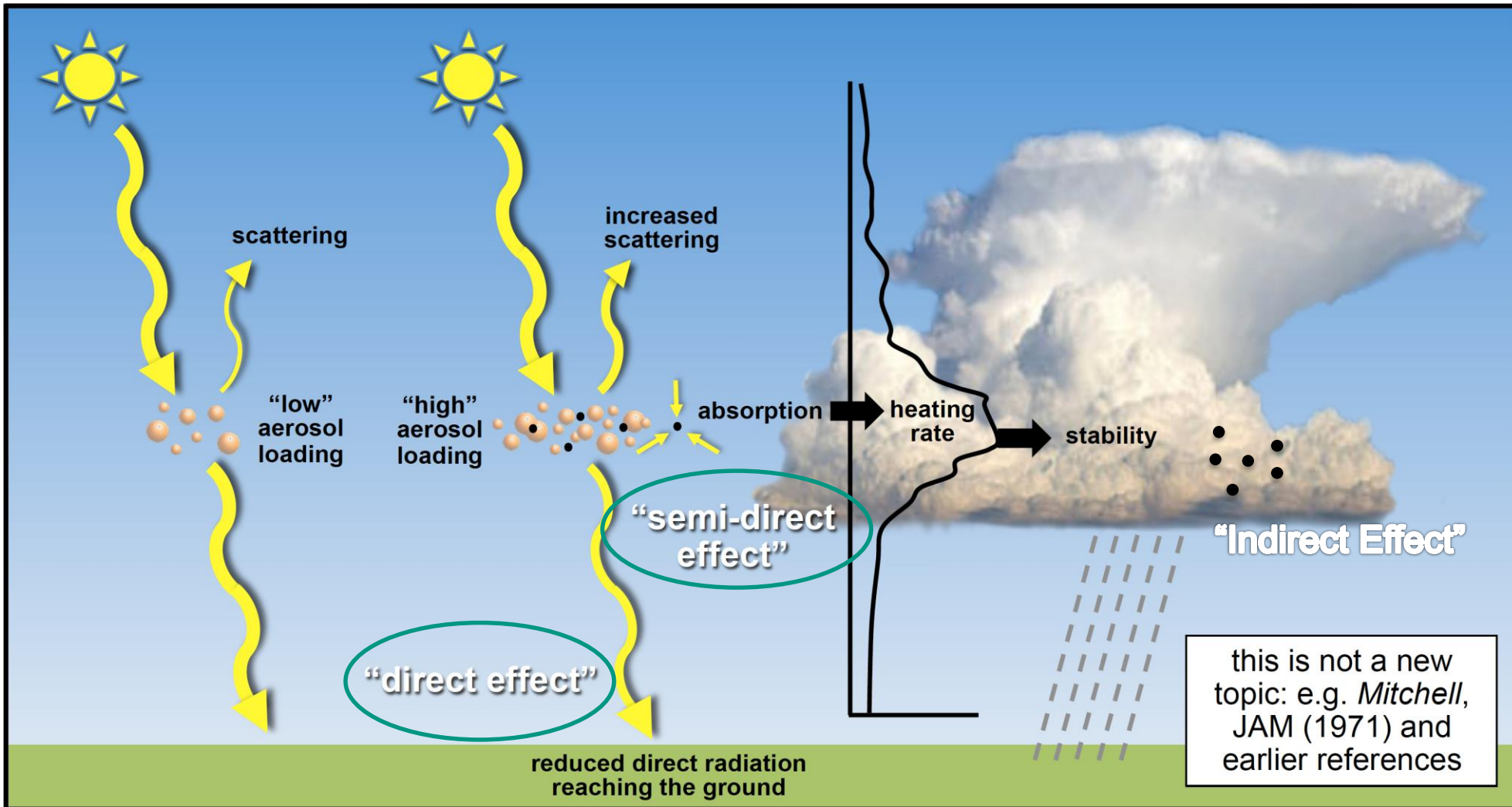




# Aerosol effects



Source: Alma Hodzic, NCAR

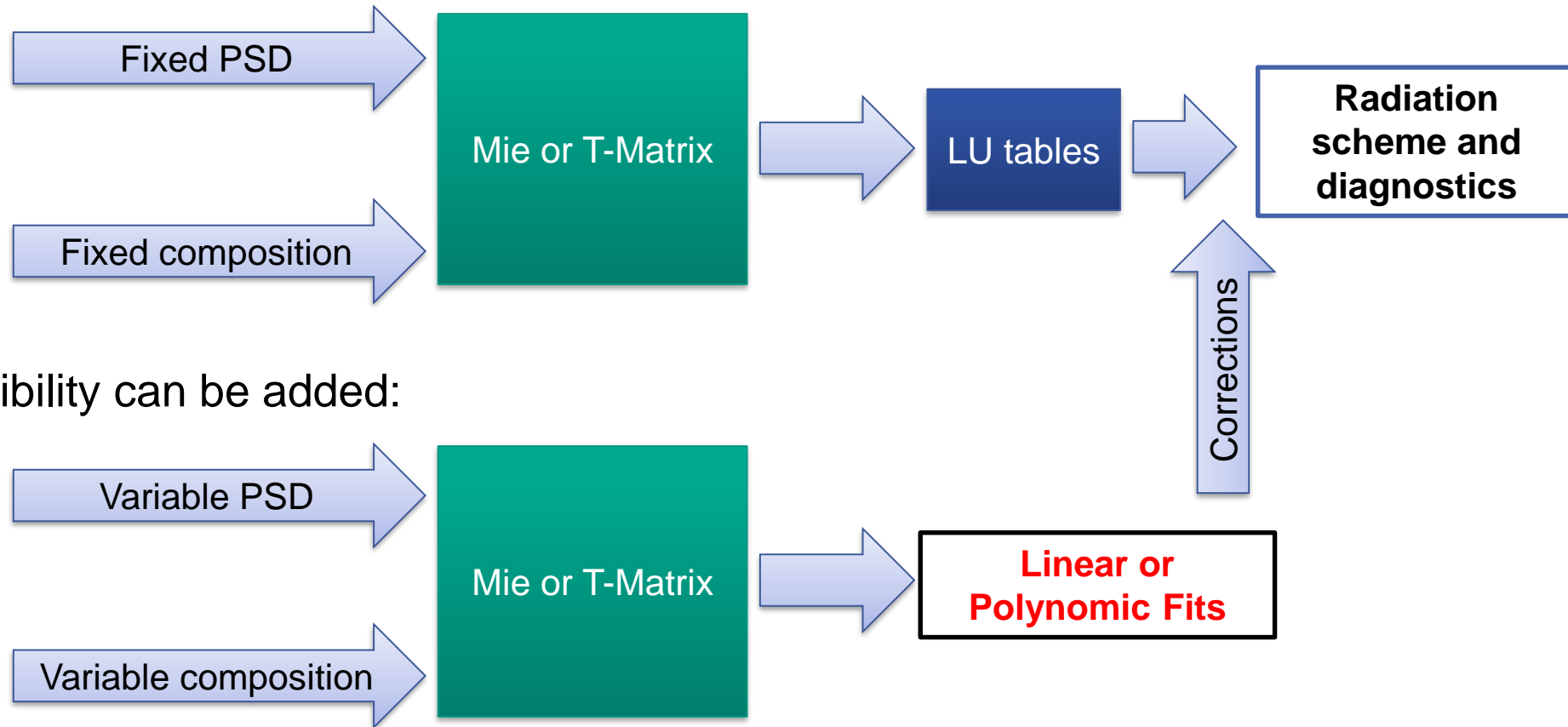
## Aerosol optical properties in models

- **Extinction Coefficient ( $Q_{ext}$ ):** fractional depletion of radiance per unit path length ( $m^{-1}$ ) due to scattering and absorption by aerosols
- **Single Scattering Albedo ( $\omega_o$ ):** ratio of scattering to extinction efficiency
- **Asymmetry Parameter ( $g$ ):** Preferred scattering direction (forward or backward) for the light encountering the aerosol particles

The values of  $Q_{ext}$ ,  $\omega$ , and  $g$  depend on

- Wavelength ( $\lambda$ ),
- Particle size distribution (PSD)
- Mixing state (externally mixed, core-shell, well-mixed ...)
- Composition (refractive index)

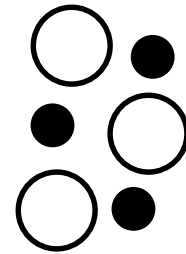
# Aerosol Optical Properties: Offline calculations



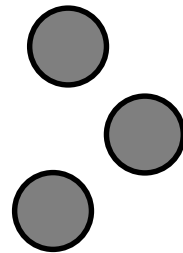
*n*-D look-up tables ... revision is necessary by changing basic assumptions.

# Effect of Aerosol Dynamics (AeroDyn)

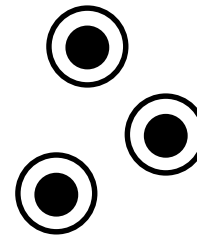
- So far: Externally mixed aerosols  $\rightarrow$  fixed RI  $\rightarrow$  lookup tables



- **New:** Internally mixed aerosols  $\rightarrow$  RI and D varies with composition  $\rightarrow$  ??

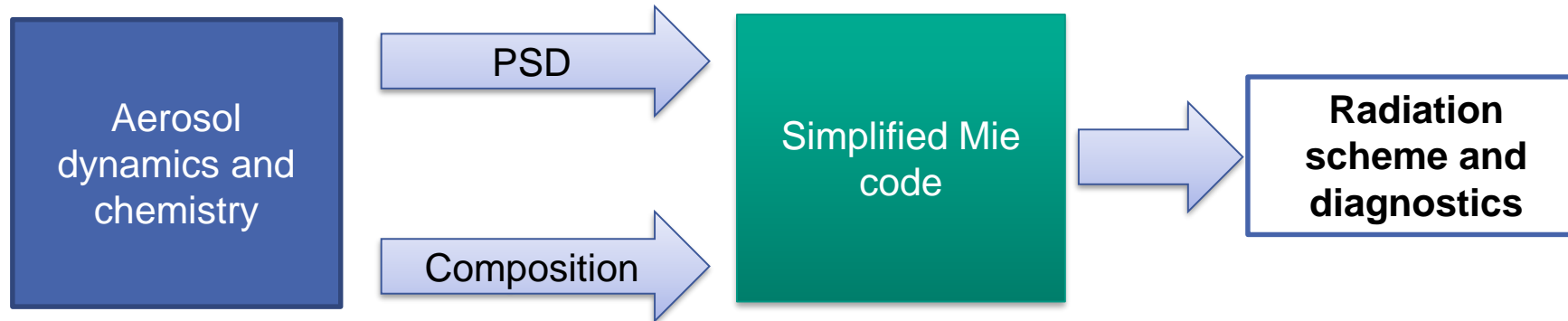


Volume-average



Core-Shell

# Aerosol Optical Properties: Online calculations

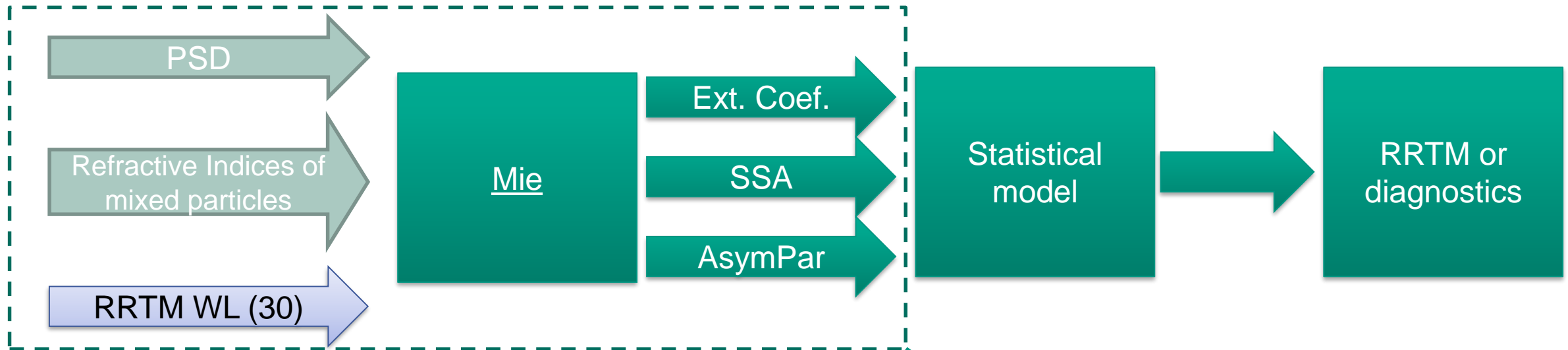


Mie codes can be computationally expensive, so an approximate version (*Ghan et al. JGR, 2001*) is also available.

Mie code is called infrequent for 3-5 WL only. For other WLs, inter- and extrapolation is necessary.

**Computationally super expensive, with low precision for inter- and extrapolated WLs ...**

## New concept



### New features:

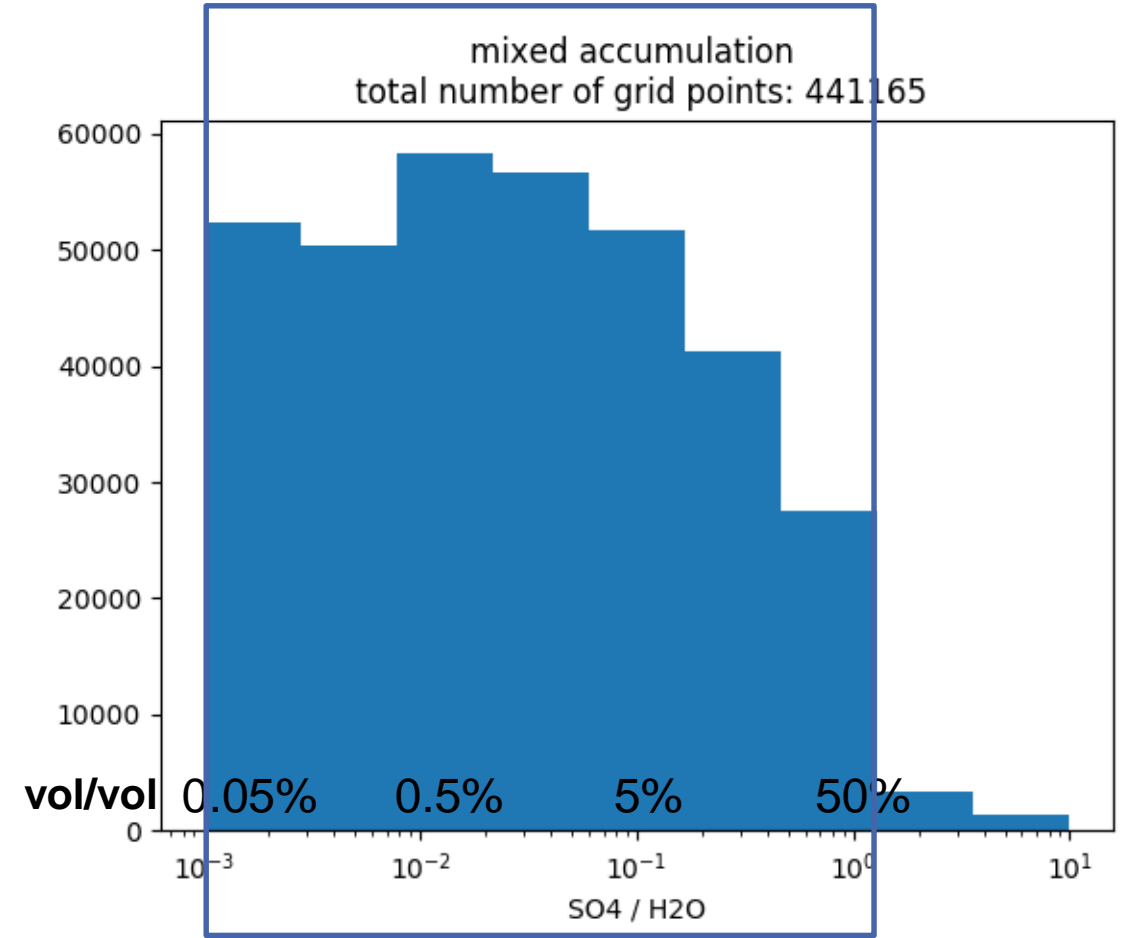
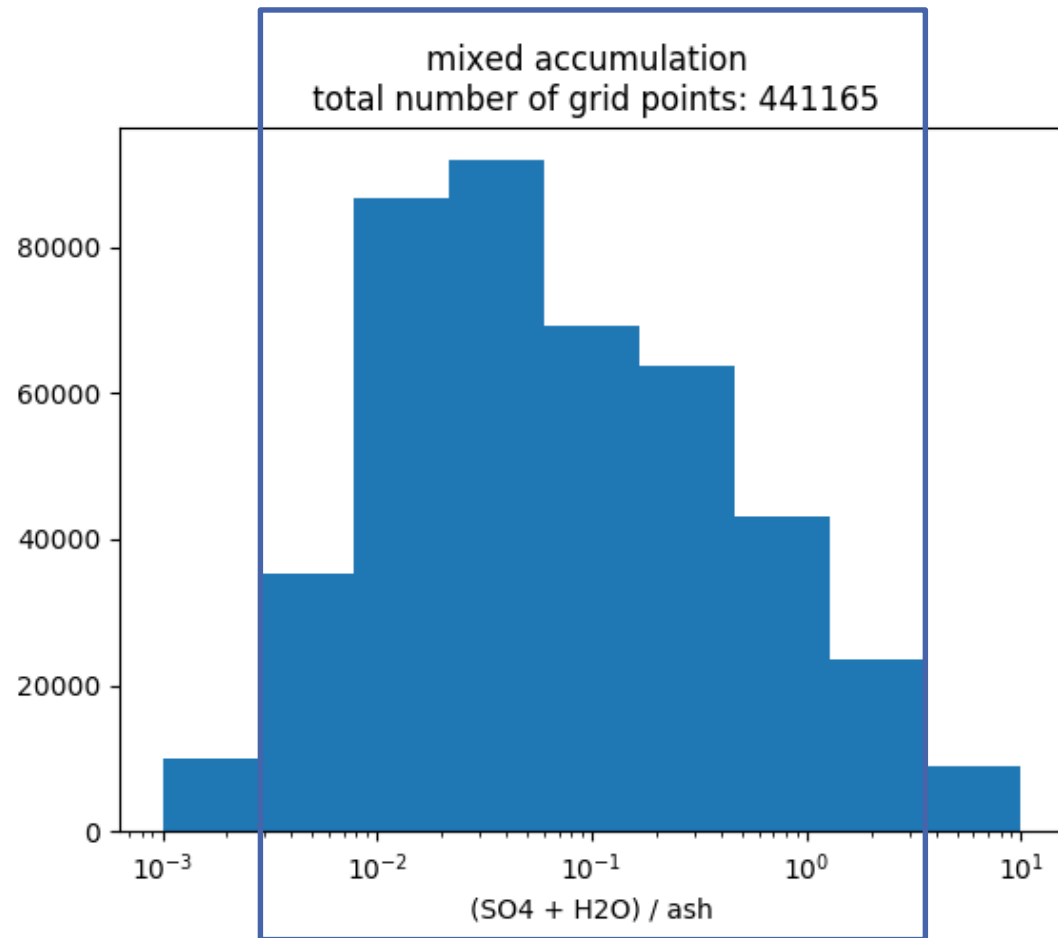
- Treatment of chemical evolution
- Treatment of PSD evolution
- Treatment of core-shell state
- One model/Fit for all

### Key challenge:

- High degrees of freedom in all above parameters
- Finding one global model/fit

**A representative and robust training dataset**

# Constraining aerosol mixing state: volcanic aerosols

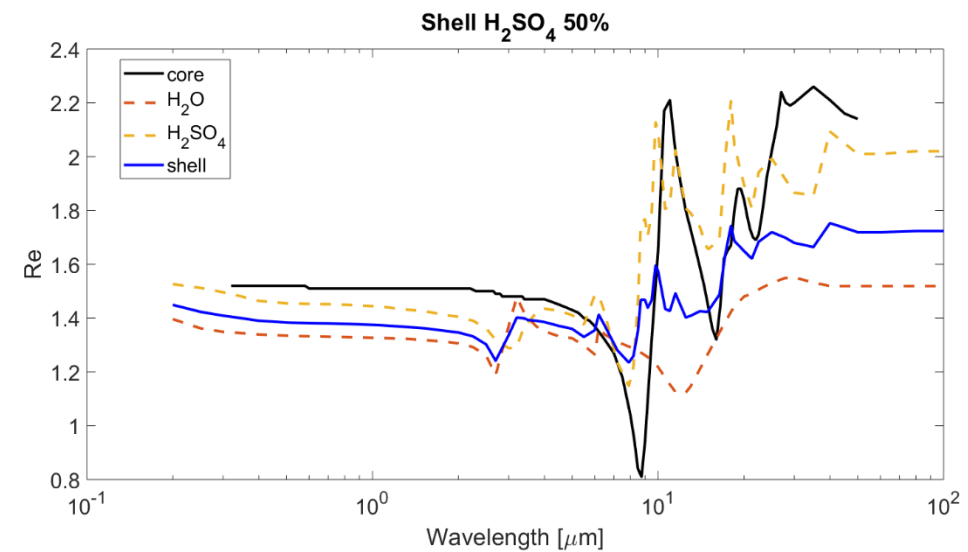
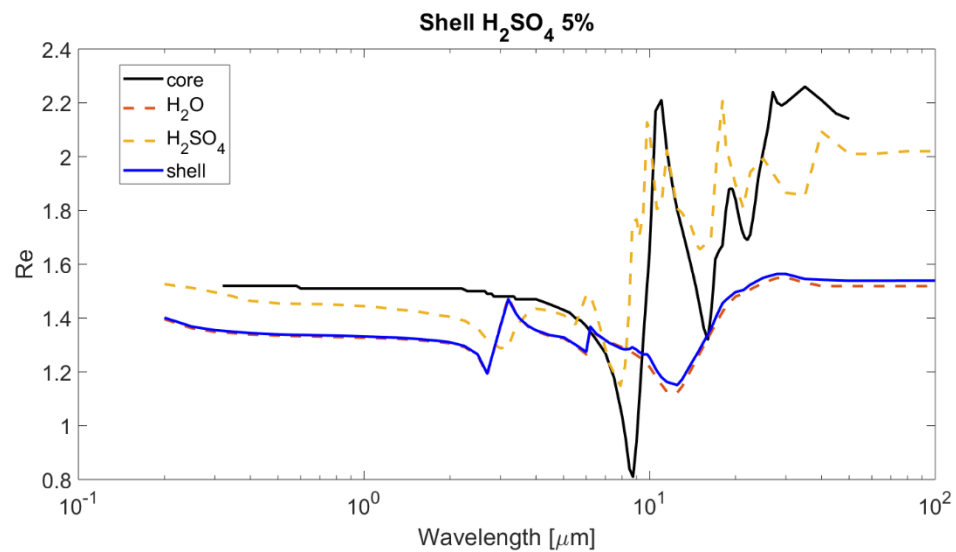
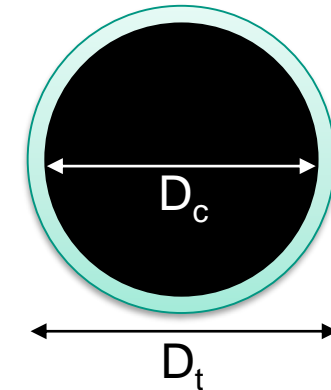




# Input variables for Mie calculations

## Variables

- $X1 = D_c/D_t$ : 0.6 to 1 (Shell thickness  $0-0.4 \cdot D_t$ )
- $X2 = D_t$ :  $0.5 \cdot D$  to  $1.5 \cdot D$  (Median diameter)
- $X3 = H_2SO_4/H_2O$ :  $1E-3$  to 1 (Sulfate concentration)
- $X4 = \lambda$ : 0.2 to 100  $\mu m$  (WL for RRTM)



# Mie calculations

To calculate the extinction and scattering factors and the asymmetry parameter:

$$Q^{\text{ext}} = \frac{2}{x^2} \sum_{n=1}^{\infty} (2n + 1) \operatorname{Re}(a_n + b_n)$$

$$Q^{\text{sca}} = \frac{2}{x^2} \sum_{n=1}^{\infty} (|a_n|^2 + |b_n|^2)$$

$$g = \frac{4}{x^2 Q^{\text{sca}}} \sum_{n=1}^{\infty} (2n + 1) \left[ \frac{n(n + 2)}{n + 1} \operatorname{Re}(a_n a_{n+1}^* + b_n b_{n+1}^*) + \frac{2n + 1}{n(n + 1)} \operatorname{Re}(a_n b_n^*) \right]$$

$$x = \frac{2 \pi r}{\lambda}$$

$$\begin{cases} a_n = \frac{\psi_n'(y) \psi_n(x) - m \psi_n(y) \psi_n'(x)}{\psi_n'(y) \xi_n(x) - m \psi_n(y) \xi_n'(x)} \\ b_n = \frac{m \psi_n'(y) \psi_n(x) - \psi_n(y) \psi_n'(x)}{m \psi_n'(y) \xi_n(x) - \psi_n(y) \xi_n'(x)} \end{cases}$$

*The extinction, scattering and absorption factors based on complex refractive indices (m), y=mx; extension for core-shell*

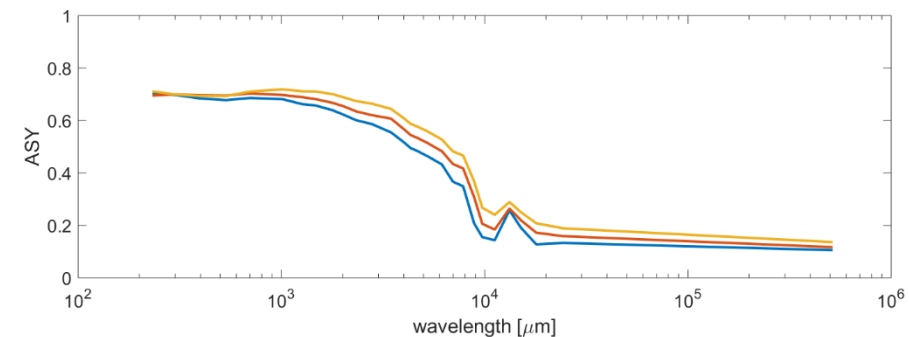
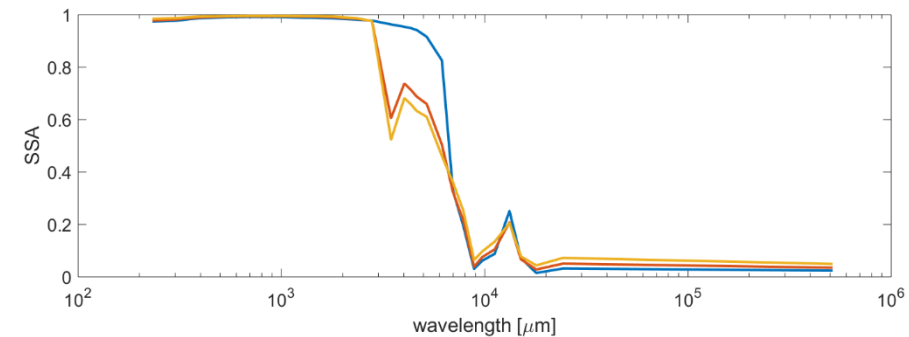
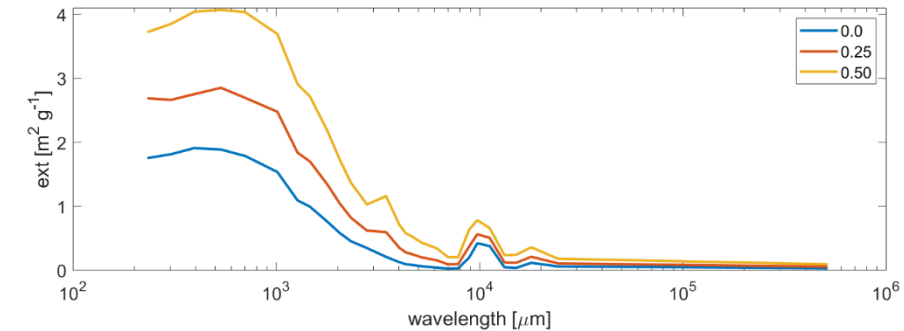
Once we have the  $Q^{\text{ext}}$ , the mass-extinction coefficient for every mode is:

$$k_e(l, \lambda, B_\lambda) = \frac{\int_0^\infty \frac{\pi}{4} d_p^2 Q_e(d_p, \lambda, B_\lambda) \psi_{0,l}(d_p) \mathrm{d}d_p}{\int_0^\infty \psi_{3,l}(d_p) \mathrm{d}d_p}$$

# Ash coated with 50% H<sub>2</sub>SO<sub>4</sub>-H<sub>2</sub>O solution (D<sub>s</sub> variable)

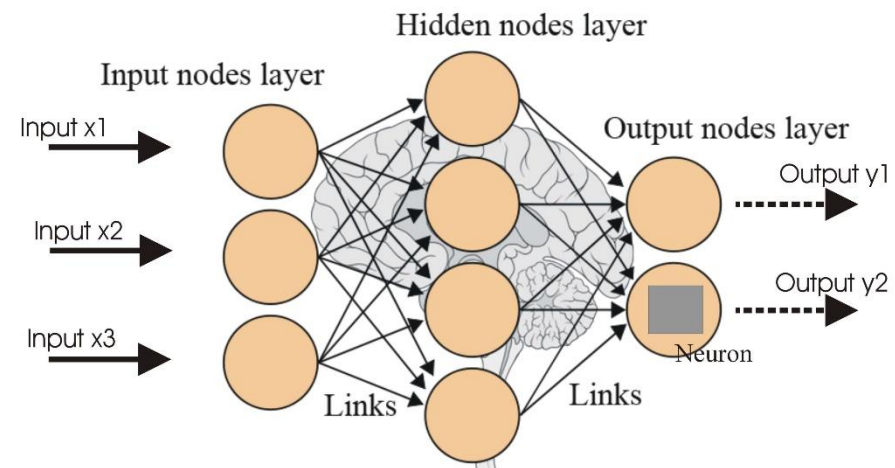
Increased coating to 0,5 of Dt leads to

- 2\* extinction at SW
- Large effect on SSA and ASY at NIR



# Statistical models

- From simple linear regression to complex artificial intelligence methods
- Chosen based on the system complexity (e.g. linear vs. nonlinear, static vs. dynamic), model performance and computational burden
- **Artificial Neural Network (ANN):** “the most recommended AI technique” as it satisfactorily learns the associations, functional dependencies and patterns in nonlinear systems resulting in excellent prediction skill.

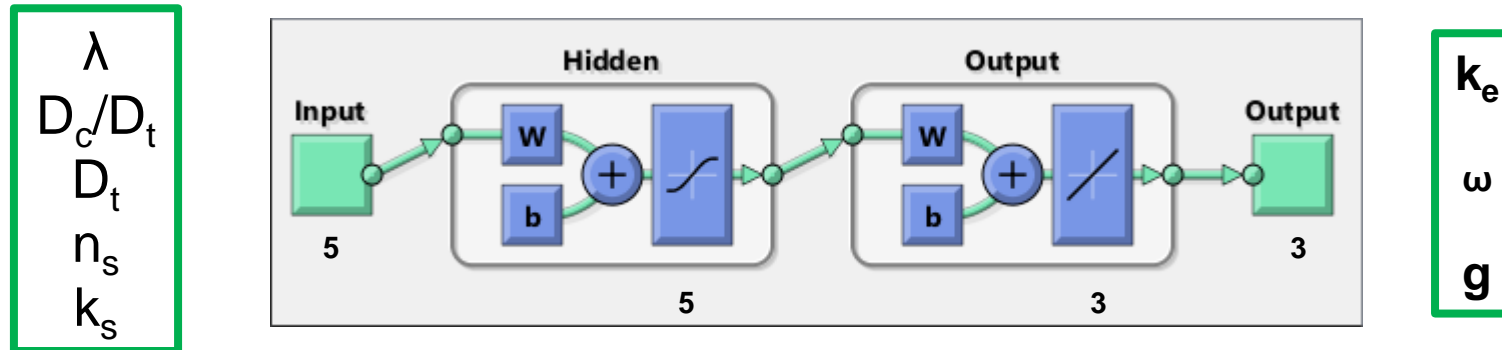


Source: <https://www.analyticsvidhya.com>

# ANN development

## ■ Network architecture

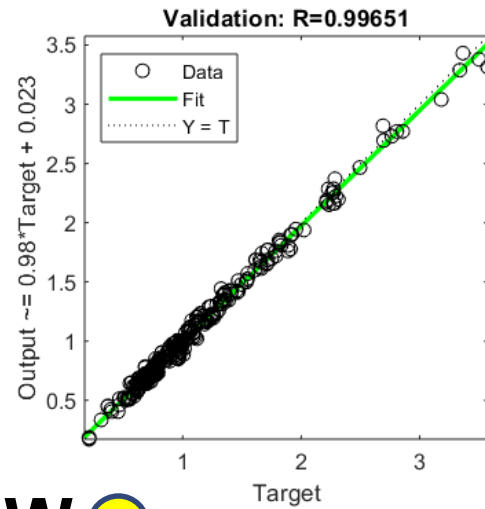
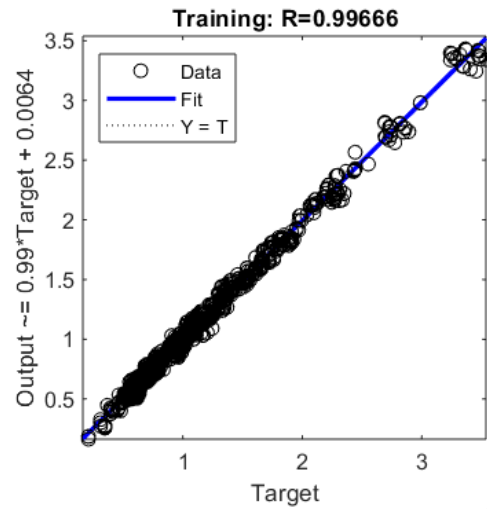
- Two-layer feed-forward network with sigmoid hidden neurons and linear output neurons



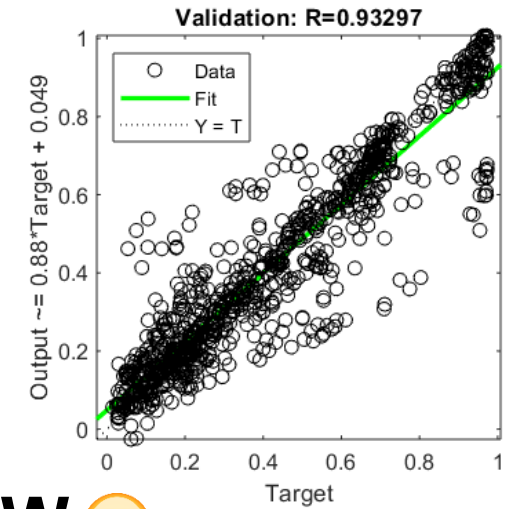
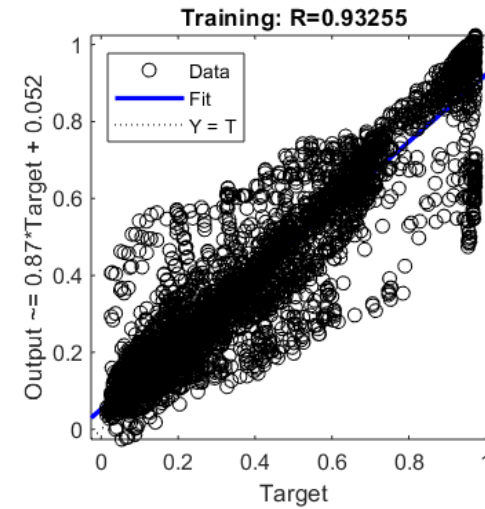
## ■ Training algorithm:

- Data division: Random (70% for Training, 15% for validation, 15% for testing)
- Training method: Levenberg-Marquardt
- Performance: mean square error
- Two networks are trained: one for SW and one for LW

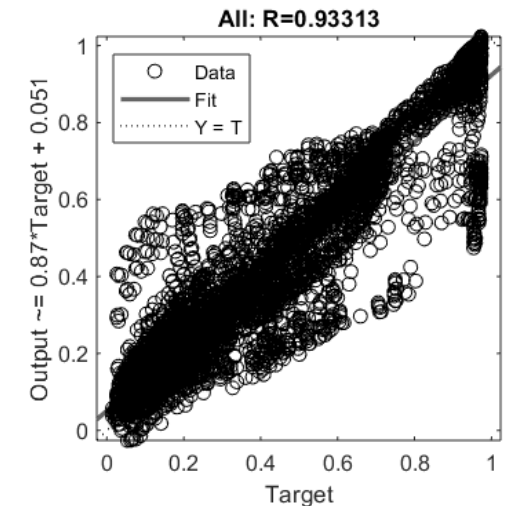
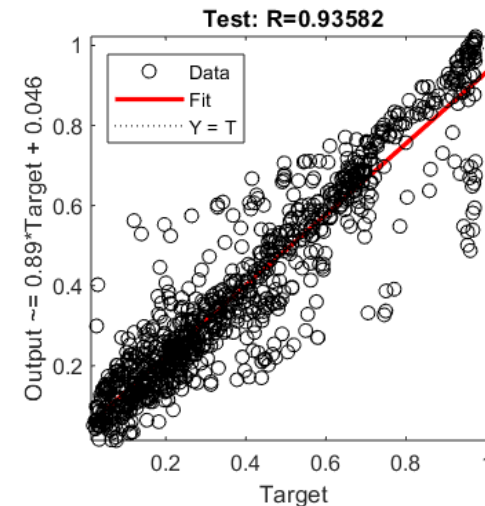
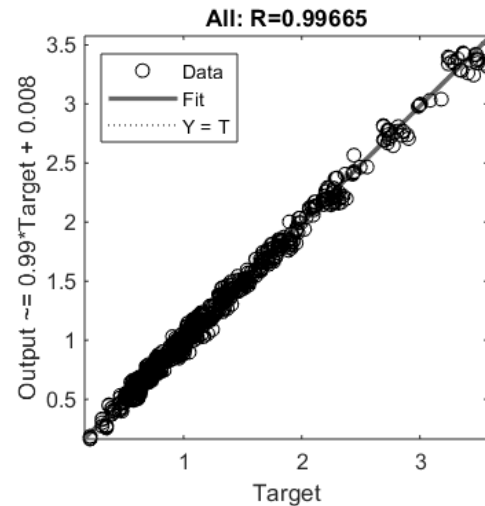
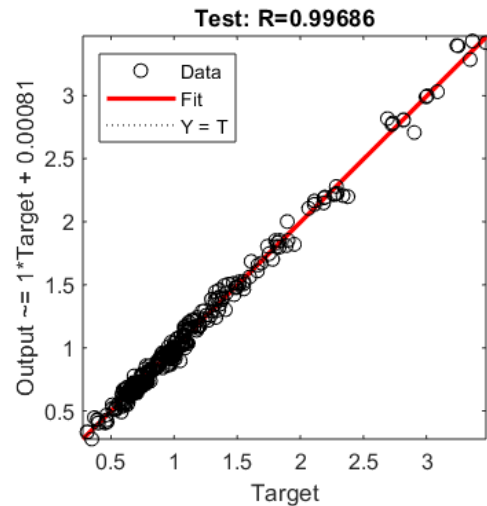
# Results of the training



**SW** 😊

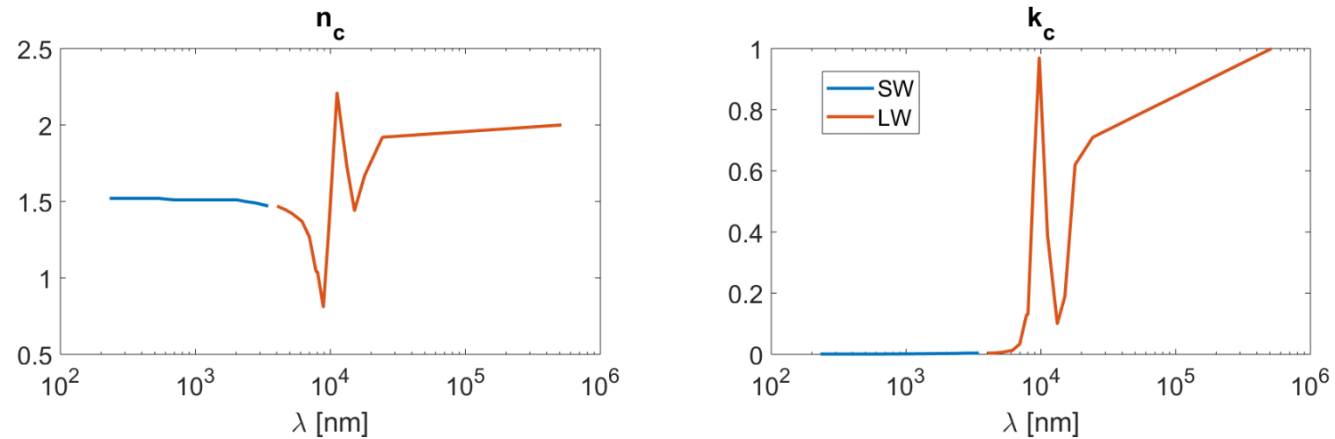


**LW** 😞

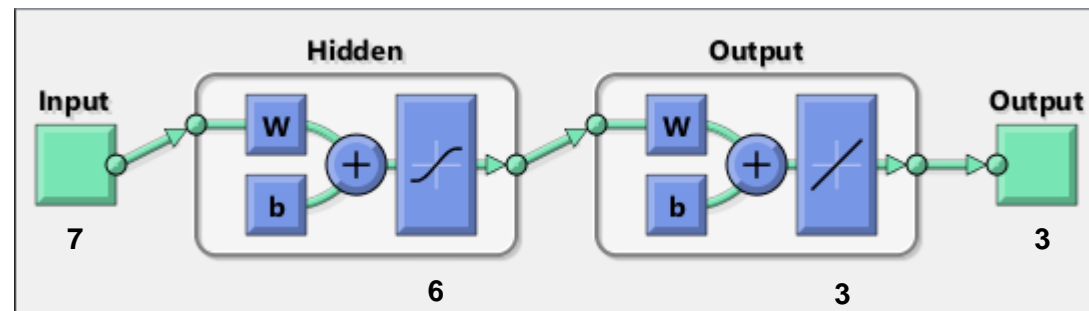


# Why is the LW problematic?

- RI of the core (ash) are highly variable in LW → it should be added to the inputs

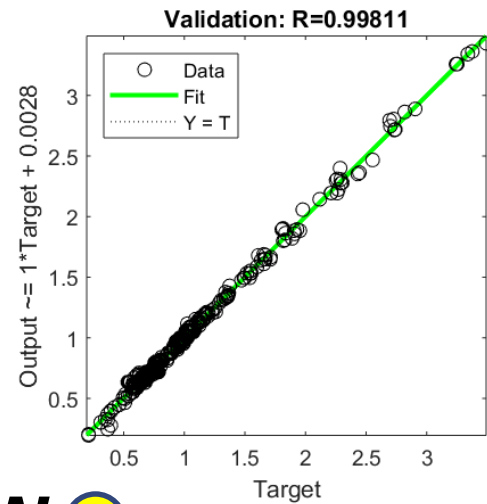
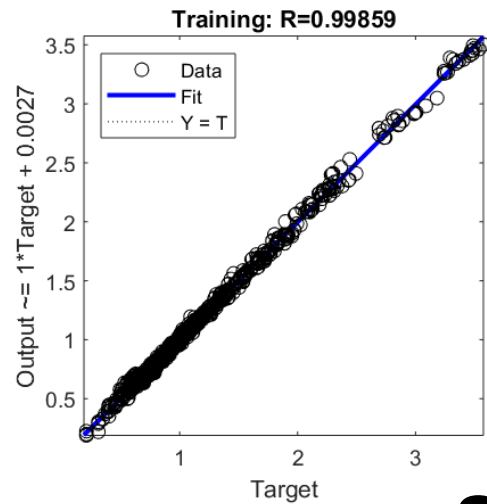


$\lambda$   
 $D_c/D_t$   
 $D_t$   
 $n_s$   
 $k_s$   
 $n_c$   
 $k_c$

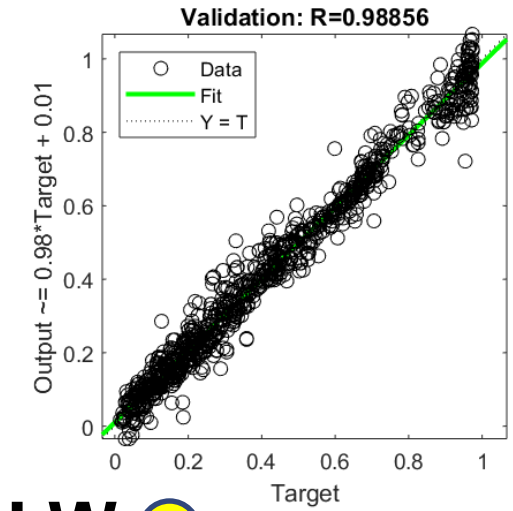
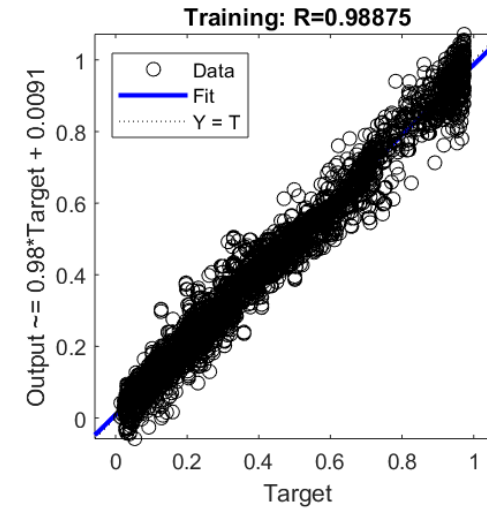


$k_e$   
 $\omega$   
 $g$

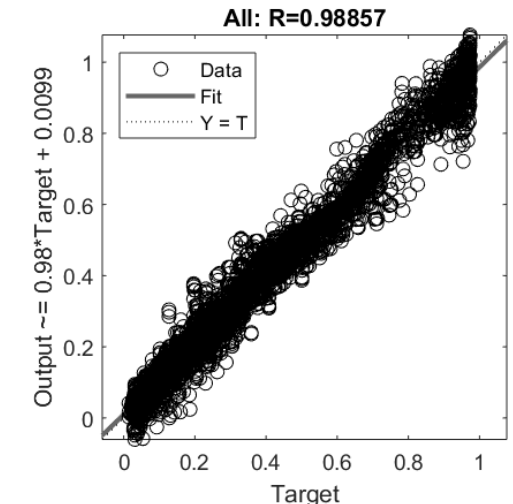
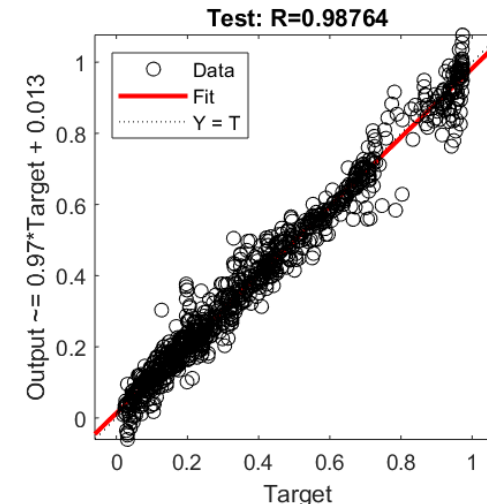
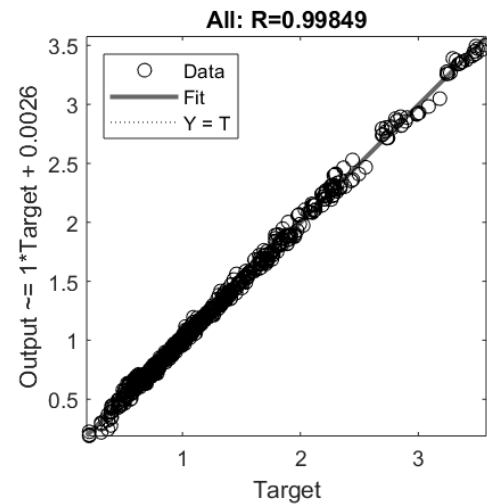
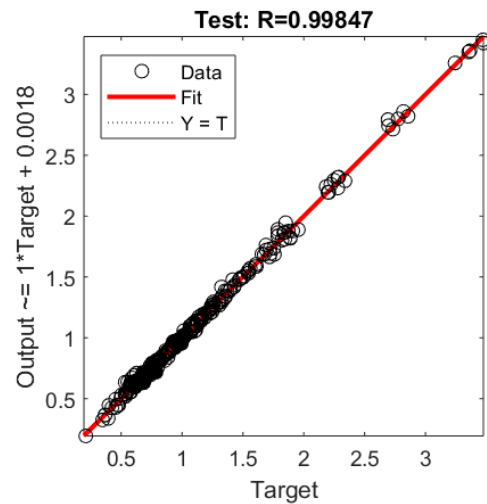
# Results of the **NEW** training



**SW** 😊

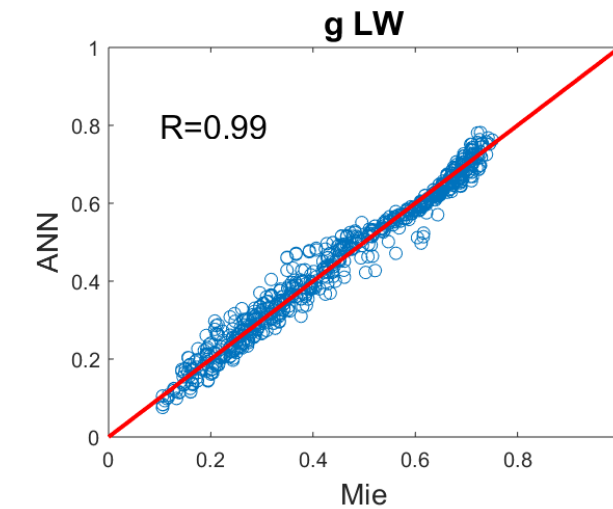
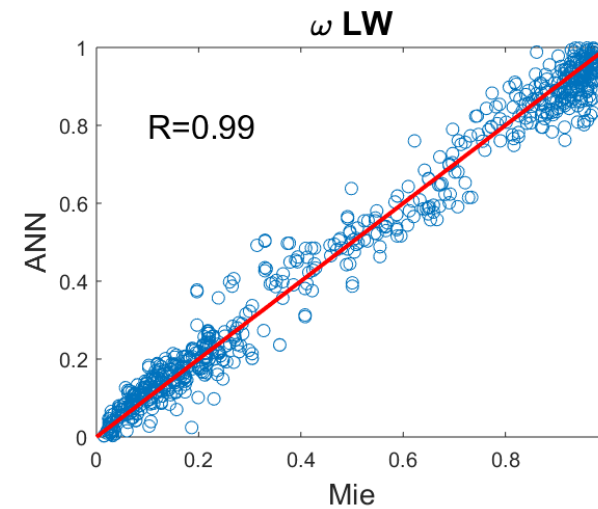
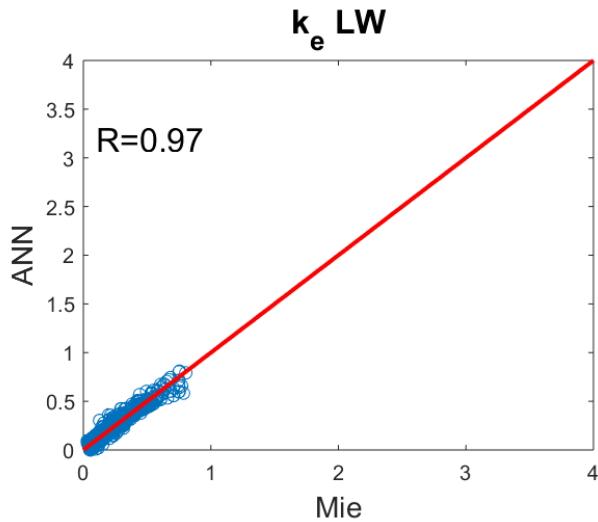
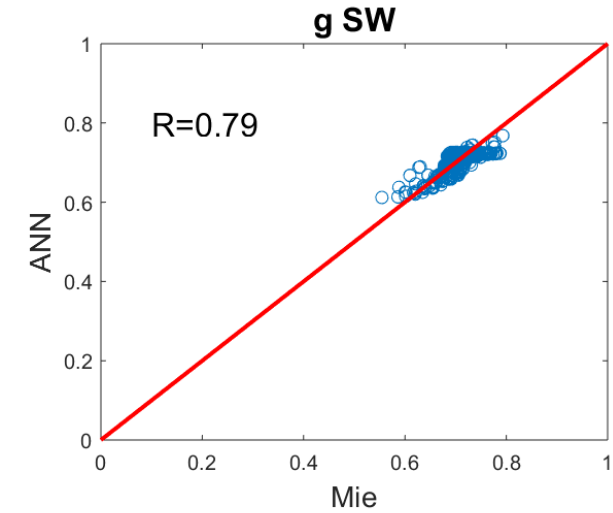
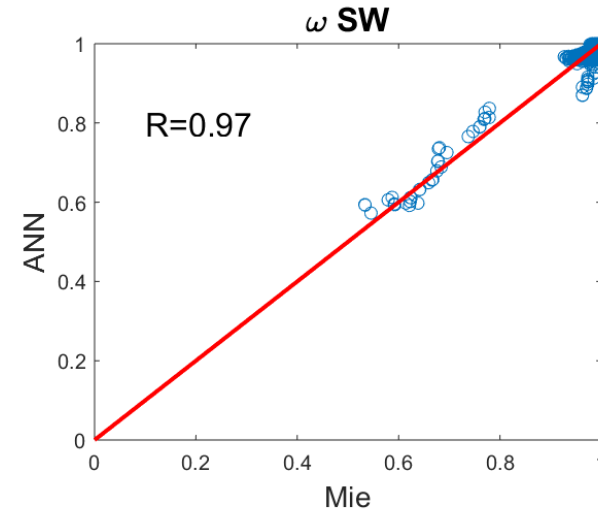
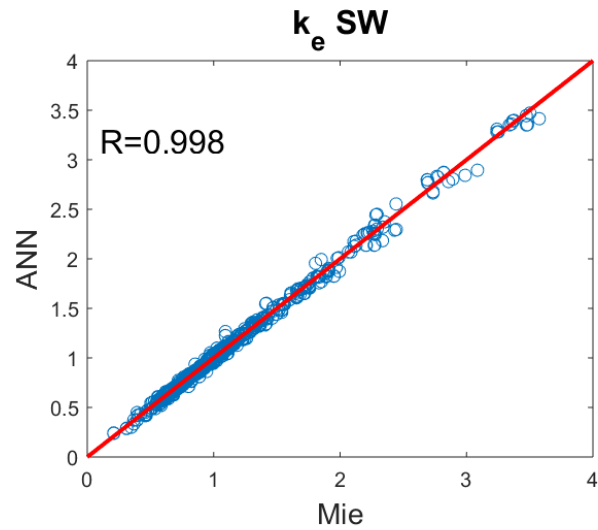


**LW** 😊





# An independent test for individual parameters



## Conclusions:

- ANN leads to  $R > 0.95$  for all parameters except  $g$  in SW with  $R \sim 0.8$
- Two ANNs for all optical parameters ...
- Proof-of-concept for a flexible, generic and computationally affordable tool for online calculation of aerosol optical properties